



Hybrid fluid-kinetic methods for space plasma simulation

Adam Stanier (Los Alamos National Laboratory)

International School/Symposium for Space Simulations (ISSS-15)
Max Planck Institute Plasma Physics, Garching, Germany

08/02/2024

LA-UR-24-27919



Managed by Triad National Security, LLC, for the U.S. Department of Energy's NNSA.

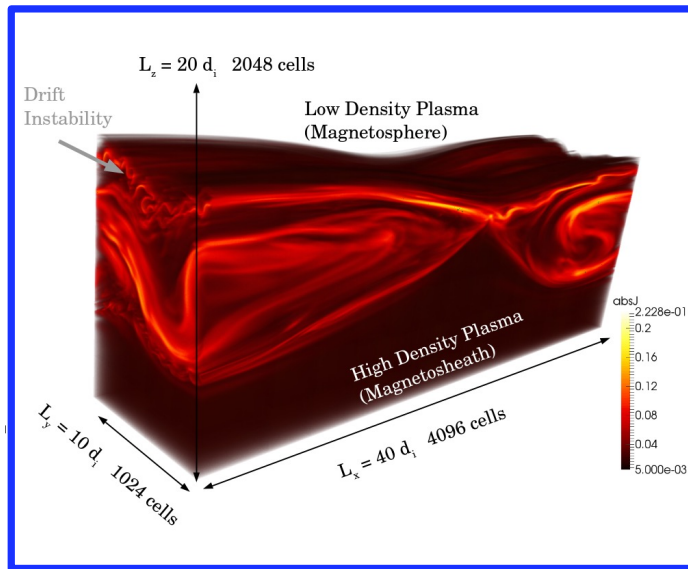
In this lecture

1. Motivating issue: scale separation in space plasma physics.
2. Introduction to hybrid fluid-kinetic methods.
3. Basic structure of a hybrid kinetic-ion fluid-electron PIC algorithm.
4. Numerical considerations to design and use a hybrid-PIC algorithm.
5. Applications of hybrid-PIC methods.
6. Fast particle kinetic hybrid model.

1. Scale separation in space plasma physics

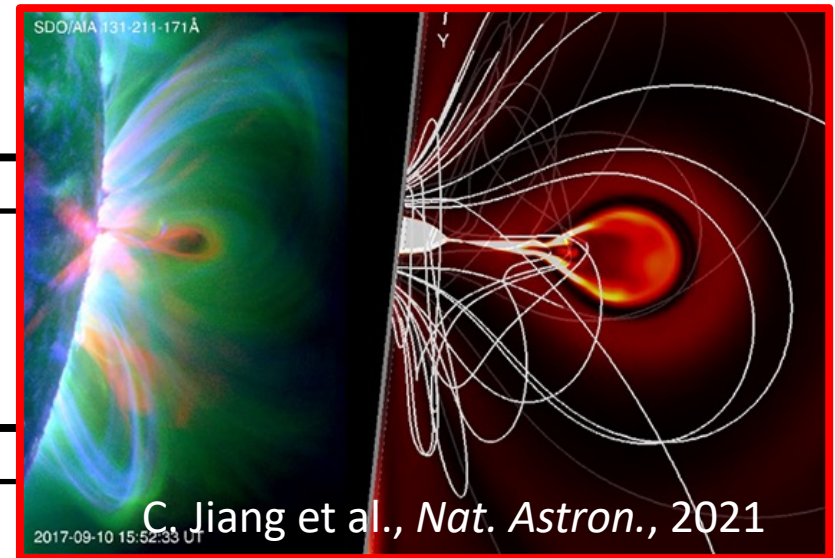
Global MHD vs local kinetic (PIC/Vlasov) modeling

Kinetic Scales



Fully Kinetic

Macroscopic Scale



MHD

Grand challenge: Modeling kinetic physics in global simulations

- **Key issues:**

1. Space plasma is virtually collisionless - invalidates MHD assumptions!
2. Kinetic PIC simulations good for studying local physics, but too expensive for global simulations.

These approaches miss potentially important coupling/feedback between different space plasma phenomena in a large global system.

E.g. shocks, reconnection, collisionless damping,...

2. Introduction to hybrid fluid-kinetic methods

Model reduction: The hybrid fluid-kinetic approach

- Choose to model:
 - Some part of the plasma kinetically (expensive).
 - Rest using a simplified fluid model (cheap).
- The choice here is problem dependent.
- Two common approaches
 - Kinetic ion + fluid electron: Suited to study coupling of large scale to ion scale behavior (trade electron scale accuracy for speed).
 - Plasma bulk fluid (MHD) + fast particle: Suited to study particle acceleration and transport by MHD-scale fluctuations.

Deriving the hybrid fluid-kinetic plasma model

1. Taking moments of the Vlasov for electrons:

Vlasov ions:

$$\partial_t f_i + \nabla \cdot (f_i \mathbf{v}) + (q_i/m_i) (\mathbf{E} + \mathbf{v} \times \mathbf{B}) \cdot \nabla_v f_i = 0.$$

Moment for electrons:

$$\partial_t n_e + \nabla \cdot (n_e \mathbf{u}_e) = 0,$$

$$m_e n_e D_t (\mathbf{u}_e) + \nabla \cdot \mathbb{P}_e + e n_e (\mathbf{E} + \mathbf{u}_e \times \mathbf{B}) = 0,$$

Pressure tensor

⋮
Infinite # equations!

Electromagnetic fields:

$$\begin{aligned} c^{-2} \partial_t \mathbf{E} &= \nabla \times \mathbf{B} - \mu_0 \mathbf{j}, & \partial_t \mathbf{B} &= -\nabla \times \mathbf{E}, \\ \nabla \cdot \mathbf{E} &= \rho/\epsilon_0, & \nabla \cdot \mathbf{B} &= 0. \end{aligned}$$

➤ So far no approximations made.. (equivalent to Vlasov-Maxwell).

Deriving hybrid model: Simplifying assumptions

- **Quasi-neutrality:** Ion and electron charge densities are approximately equal:

$$\rho/en_e = (q_i n_i - e n_e)/en_e \sim 0.$$

➤ Hybrid model breaks down in vacuum regions!

- Non-relativistic ($c \rightarrow \infty$, $\epsilon_0 \rightarrow 0$): Removes light waves from the model.

➤ These modify the Maxwell equations as:

1. Gauss' law:
$$\nabla \cdot \mathbf{E} = \frac{q_i n_i - e n_e}{\epsilon_0} \rightarrow \frac{0}{0}!$$

2. Maxwell-Ampere:
$$c^{-2} \partial_t \mathbf{E} = \nabla \times \mathbf{B} - \mu_0 \mathbf{j}$$

- How to we calculate electric field now? **Electron momentum equation!**

Deriving hybrid model: Simplifying assumptions

- Full Ohm's law:

$$\mathbf{E} = \underbrace{-\mathbf{u}_e \times \mathbf{B}}_{\text{Convective}} - \frac{m_e D_t \mathbf{u}_e}{e} \overset{\text{Inertial}}{\quad} - \frac{\nabla \cdot \mathbb{P}_e}{en} \overset{\text{Pressure tensor}}{\quad},$$

$$- \mathbf{u}_i \times \mathbf{B} + \frac{\mathbf{j} \times \mathbf{B}}{en} \overset{\text{Hall}}{\quad}$$
- At this point still have infinite moment equations (\mathbb{P}_e , etc): need closure!
- Many possible choices to truncate & simplify.
- Typically, we choose simply:
 - Isothermal electrons: $\mathbb{P}_e \sim T_{e0} n$,
 - Adiabatic electrons: $\mathbb{P}_e \sim T_{e0} n^\gamma$ with $\gamma=5/3$.
- Another common simplification is $m_e \rightarrow 0$ (remove electron kinetic scales).

Hybrid full-orbit ion and fluid electron model

- With all of the above assumptions we have:

Full-orbit ions: $\partial_t f_i + \nabla \cdot (f_i \mathbf{v}) + (q_i/m_i) (\mathbf{E} + \mathbf{v} \times \mathbf{B}) \cdot \nabla_v f_i = 0.$

Ohm's law: $\mathbf{E} = -\mathbf{u}_i \times \mathbf{B} + \frac{\mathbf{j} \times \mathbf{B} - \nabla p_e}{en},$

Faraday: $\partial_t \mathbf{B} = -\nabla \times \mathbf{E},$

Ampere: $\mathbf{j} = (\nabla \times \mathbf{B}) / \mu_0,$

Solenoidal: $\nabla \cdot \mathbf{B} = 0.$

Adiabatic pressure: $p_e = T_{e0} n_0 (n/n_0)^\gamma.$

Closed via:

Quasi-neutrality:

$$n = n_i = \frac{1}{e} \int q_i f_i d^3 v,$$

Ion current carrying velocity:

$$\mathbf{u}_i = \frac{1}{en} \int q_s \mathbf{v} f_i d^3 v.$$

Normalization: Ion units for EM hybrid model

- Take some reference parameters:
 - B_0 magnetic field, n_0 density, m_0 ion mass, v_0 velocity, L_0 distance, T_0 temperature.

- Write each variable $\chi = \chi_0 \hat{\chi}$
 - χ_0 Reference value
 - $\hat{\chi}$ Normalized variable

➤ Ohm's law:
$$\hat{\mathbf{E}} = -\hat{\mathbf{u}}_i \times \hat{\mathbf{B}} + \frac{\hat{d}_i}{\hat{n}} \left((\hat{\nabla} \times \hat{\mathbf{B}}) \times \hat{\mathbf{B}} - \hat{\nabla} \hat{p}_e \right)$$

- Here $\hat{d}_i = \frac{d_i}{L_0}$,

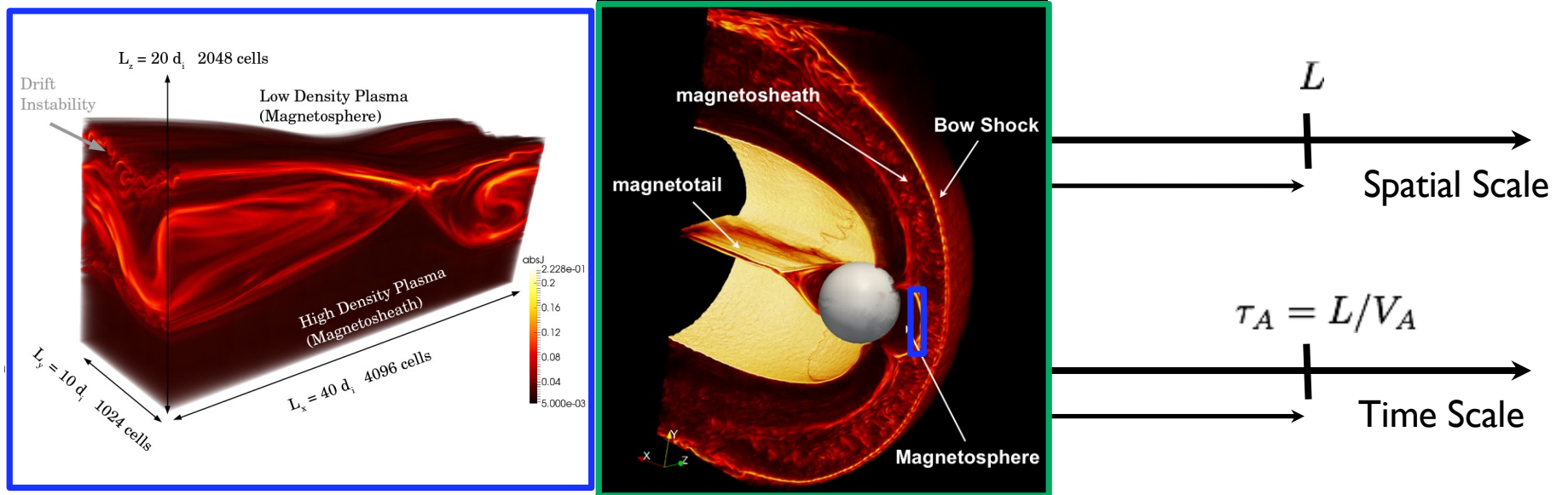
$d_i = \frac{v_A}{\Omega_{ci}}$	$v_A = \frac{B_0}{\sqrt{\mu_0 n_0 m_0}}$	$\Omega_{ci}^{-1} = \frac{m_0}{e B_0}$
Ion skin-depth (length)	Alfven speed (velocity)	Inverse gyro-freq (time)

- Drop '^' notation for normalized units. In these ion units $\hat{d}_i = 1$.

Scale bridging via hybrid kinetic-ion fluid-electron model

Kinetic Scales

Macroscopic Scale



Fully Kinetic

MHD

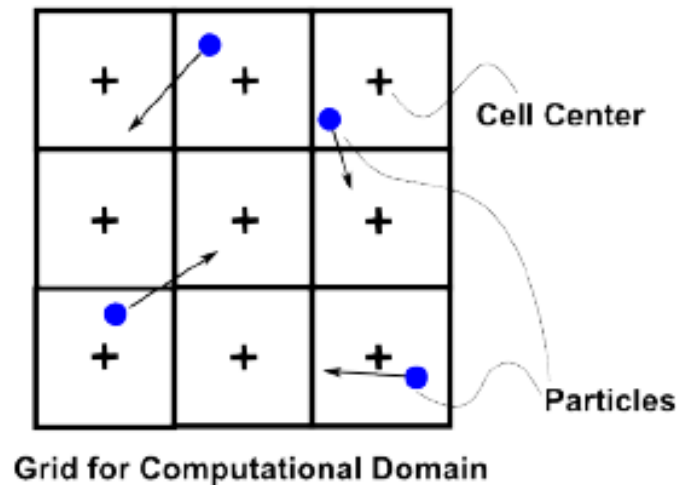
Hybrid
(or Gyrokinetic, or Multi-Fluid)

Methods to solve the hybrid model equations

- Two main approaches to solve the hybrid model.
 1. Continuum:
 - Discretize the 6D distribution function $f_i(\mathbf{x}, \mathbf{v})$ on a spatial mesh.
 - Noise free, but can be have a large memory footprint (curse of dimensionality).
 2. Hybrid particle-in-cell:
 - Ion distribution $f_i(\mathbf{x}, \mathbf{v})$ represented by particles.
 - Subject to finite particle noise $\sim 1/\sqrt{N_p}$.
- I will focus on this second approach in this lecture:
 - Particle-In-Cell solution to the kinetic-ion fluid-electron plasma model.

3. Basic structure of a hybrid-PIC algorithm

Kinetic ions are discretized with marker particles



- Vlasov equation:

$$\partial_t f_i + \nabla \cdot (f_i \mathbf{v}) + (q_i/m_i) (\mathbf{E} + \mathbf{v} \times \mathbf{B}) \cdot \nabla_v f_i = 0.$$

- Discrete sum of marker particles:

$$f_i = \sum_p S[\mathbf{x} - \mathbf{x}_p(t)] \delta[\mathbf{v} - \mathbf{v}_p(t)]$$

Finite-size in real space

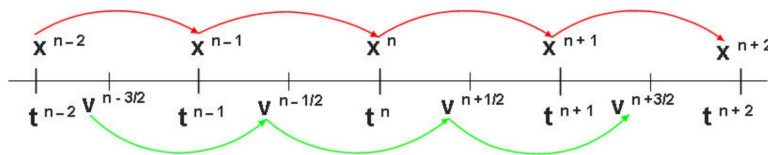
Point-like in velocity space



$$\begin{cases} \frac{d\mathbf{x}_p}{dt} = \mathbf{v}_p, \\ \frac{d\mathbf{v}_p}{dt} = \frac{q_i}{m_i} (\mathbf{E}_p + \mathbf{v}_p \times \mathbf{B}_p). \end{cases}$$

Particle equations of motion: Time discretization

- Ions may gyrate >100-1000 times around the magnetic field within a simulation.
- Key requirements:
 - Long term orbit accuracy,
 - Preserve kinetic energy (gyroradius).
- Most used are variants of the *leapfrog* method.

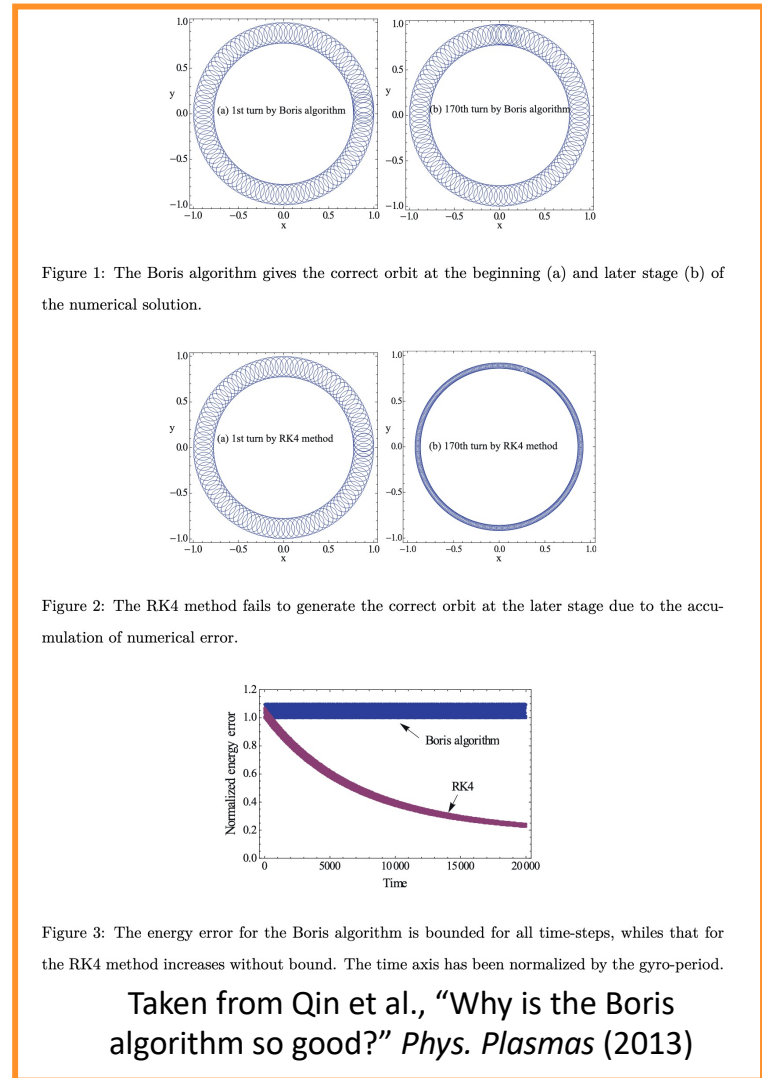


- Boris-push for velocity update:

Half-step with E:
$$\mathbf{v}_p^- = \mathbf{v}_p^{n-1/2} + \frac{\Delta t}{2} \frac{q_p}{m_p} \mathbf{E}_p^n,$$

Rotation about B:
$$\mathbf{v}_p^+ - \mathbf{v}_p^- = \Delta t \frac{q_p}{m_p} \left(\frac{\mathbf{v}_p^+ + \mathbf{v}_p^-}{2} \right) \times \mathbf{B}_p^n,$$

Half-step with E:
$$\mathbf{v}_p^{n+1/2} = \mathbf{v}_p^+ + \frac{\Delta t}{2} \frac{q_p}{m_p} \mathbf{E}_p^n.$$

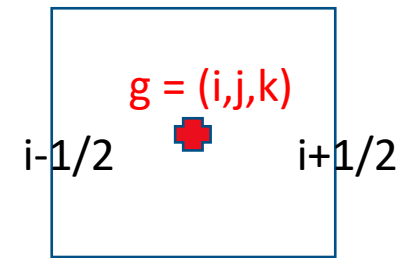


Mesh-based fluid electrons/EM fields

- Unlike typical MHD & PIC codes, most hybrid-PIC codes use a very simple cell-centered discretization:

$$\mathbf{E}_g = -\mathbf{u}_g \times \mathbf{B}_g + \frac{(\nabla \times \mathbf{B}_g) \times \mathbf{B}_g - \nabla(p_e)_g}{en_g} = 0,$$

$$\partial_t \mathbf{B}_g = -\nabla \times \mathbf{E}_g.$$



Where e.g. $\nabla \chi_g = \frac{\chi_{i+1,j,k} - \chi_{i-1,j,k}}{2\Delta x} \hat{\mathbf{x}} + \dots$

- Centered discretization of advection terms can be unstable!
 - But in hybrid advection is handled by particles!
- Cell centered does conserve $\nabla \cdot \mathbf{B}_g = 0$.
- Explicit time-stepping for hybrid is much more complicated! – in a few slides.

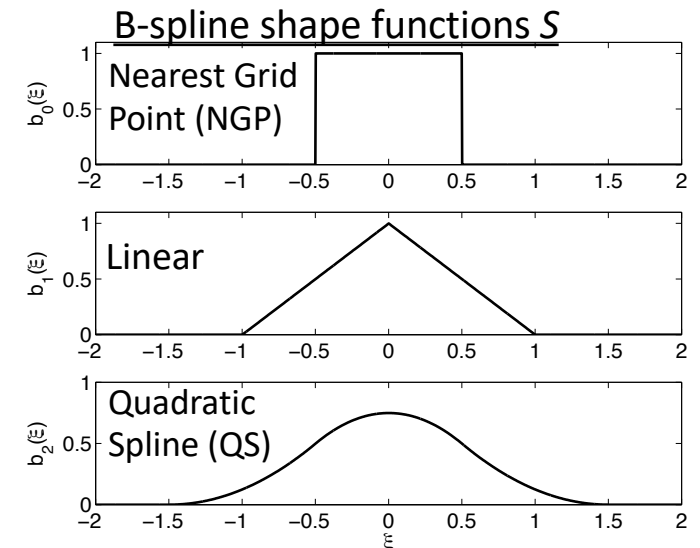
Particle-Mesh interpolation

- Scatter fields: $\mathbf{E}_p = \sum_g S(\mathbf{x}_g - \mathbf{x}_p) \mathbf{E}_g,$

$$\mathbf{B}_p = \sum_g S(\mathbf{x}_g - \mathbf{x}_p) \mathbf{B}_g$$

- Gather moments: $n_g = \frac{1}{e} \sum_p q_p S(\mathbf{x}_g - \mathbf{x}_p) / \Delta V,$

$$\mathbf{u}_g = \frac{1}{en_g} \sum_p q_p \mathbf{v}_p S(\mathbf{x}_g - \mathbf{x}_p) / \Delta V$$



- PIC codes often spend most runtime doing these particle-mesh interpolations!
- Higher order shape functions reduce noise, but more expensive.
- Extra care needs to be used with higher order shape functions in hybrid-PIC due to the *hybrid cancellation problem!* (see later).

Additional spatial smoothing

- In my experience, hybrid-PIC seems to be more noisy than regular Vlasov-Maxwell PIC.
- No proof of this statement at present, but consider:

$$p_e = T_{e0} n_0 (n/n_0)^\gamma. \quad \mathbf{E} = -\mathbf{u}_i \times \mathbf{B} + \frac{\mathbf{j} \times \mathbf{B} - \nabla p_e}{en},$$

Noisy (from particles)
Gradient of noisy quantity amplifies noise

- Smoothing often used in hybrid-PIC.

Quick and effective:

Binomial filter 'SM': $\frac{1}{4} \begin{bmatrix} 1 \\ 2 \\ 1 \end{bmatrix} * \frac{1}{4} \begin{bmatrix} 1 & 2 & 1 \end{bmatrix} \longrightarrow \frac{1}{16} \begin{bmatrix} 1 & 2 & 1 \\ 2 & 4 & 2 \\ 1 & 2 & 1 \end{bmatrix}$

Smooth moments: $n_g \rightarrow \text{SM}(n_g), \quad \mathbf{u}_g \rightarrow \text{SM}((n\mathbf{u})_g) / \text{SM}(n_g),$

& smooth fields: $\mathbf{E}_p^* = \sum_g S(\mathbf{x}_g - \mathbf{x}_p) \text{SM}(\mathbf{E}_g^*), \quad \mathbf{B}_p = \sum_g S(\mathbf{x}_g - \mathbf{x}_p) \text{SM}(\mathbf{B}_g).$

4. Numerical considerations to design and use a hybrid-PIC algorithm

To compare: Explicit time-stepping in Vlasov-Maxwell PIC

- In Vlasov-Maxwell PIC, explicit time-stepping fits together well:

- Assume we know: $\mathbf{v}_p^{n-1/2}, \mathbf{x}_p^n, \mathbf{B}^n, \mathbf{E}^n$
↑
↑
Staggered to leapfrog

Leap-frog particles
(second order,
Volume preserving)

$$\left\{ \begin{array}{l} \mathbf{v}_p^{n+1/2} = \mathbf{v}_p^{n-1/2} + \Delta t \left[\mathbf{E}_p^n + \frac{(\mathbf{v}_p^{n+1/2} + \mathbf{v}_p^{n-1/2})}{2} \times \mathbf{B}^n \right], \\ \mathbf{x}_p^{n+1} = \mathbf{x}_p^n + \Delta t \mathbf{v}_p^{n+1/2}. \end{array} \right.$$

Verlet fields
(leapfrog)

$$\left\{ \begin{array}{l} \mathbf{B}^{n+1/2} = \mathbf{B}^n - (\Delta t/2) \nabla \times \mathbf{E}^n, \\ \mathbf{E}^{n+1} = \mathbf{E}^n + \Delta t \left[\nabla \times \mathbf{B}^{n+1/2} - \mathbf{j}^{n+1/2} \right], \\ \mathbf{B}^{n+1} = \mathbf{B}^{n+1/2} - (\Delta t/2) \nabla \times \mathbf{E}^{n+1}. \end{array} \right.$$

Can we take the same approach in hybrid-PIC?

- Assume known $(\mathbf{v}_p^{n-1/2}, \mathbf{x}_p^n, \mathbf{B}^n, \mathbf{E}^n)$, then

Explicit Boris push
(second order,
Volume preserving)

$$\left\{ \begin{array}{l} \mathbf{v}_p^{n+1/2} = \mathbf{v}_p^{n-1/2} + \Delta t \left[\mathbf{E}_p^n + \frac{(\mathbf{v}_p^{n+1/2} + \mathbf{v}_p^{n-1/2})}{2} \times \mathbf{B}^n \right], \\ \mathbf{x}_p^{n+1} = \mathbf{x}_p^n + \Delta t \mathbf{v}_p^{n+1/2}. \end{array} \right.$$

Static Ohm's law:

$$\mathbf{B}^{n+1/2} = \mathbf{B}^n - (\Delta t/2) \nabla \times \mathbf{E}^n,$$

$$\mathbf{E}^{n+1/2} = -\mathbf{u}^{n+1/2} \times \mathbf{B}^{n+1/2} + \dots$$

$$\mathbf{B}^{n+1} = \mathbf{B}^n - \Delta t \nabla \times \mathbf{E}^{n+1/2}.$$

Explicit midpoint
(second order)

Lastly:

$$\mathbf{E}^{n+1} = -\mathbf{u}^{n+1} \times \mathbf{B}^{n+1} + \dots$$

Unknown! We have **implicit coupling** $\mathbf{E}^{n+1} = \mathbf{E}^{n+1} \left(\mathbf{v}_p^{n+3/2} (\mathbf{E}^{n+1}) \right)$.

Explicit hybrid-PIC time-stepping schemes

- This has been dealt with (explicitly) via:

- Velocity moment extrapolation:

$$\mathbf{u}^{n+1} = \frac{3}{2}\mathbf{u}^{n+1/2} - \frac{1}{2}\mathbf{u}^{n-1/2}.$$

- Predictor-corrector methods:

Predict:
$$\begin{cases} \mathbf{E}'^{n+1} = 2\mathbf{E}^{n+1/2} - \mathbf{E}^n, \\ \mathbf{E}'^{n+1} \xrightarrow{\text{push}} v_p'^{n+3/2} \xrightarrow{\text{Faraday/Ohm}} (\mathbf{B}, \mathbf{E}')^{n+3/2}, \end{cases}$$

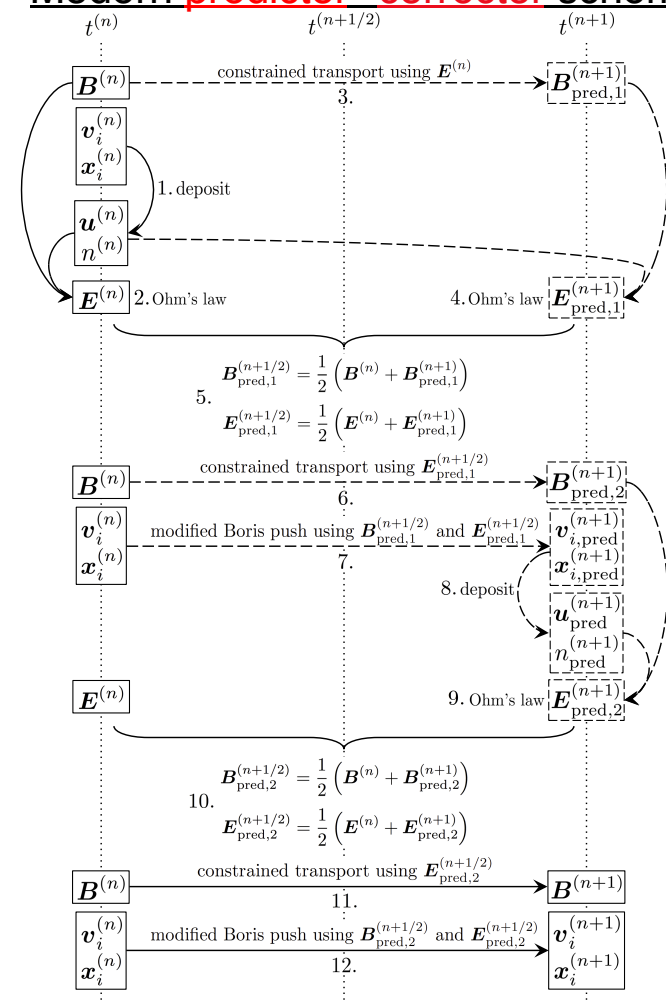
Correct:
$$\mathbf{E}^{n+1} = \frac{1}{2} \left(\mathbf{E}'^{n+3/2} + \mathbf{E}^{n+1/2} \right).$$

- Moment methods:

$$(n\mathbf{u})^{n+1} = (n\mathbf{u})^{n+1/2} + \frac{e\Delta t}{2m} \left[n^{n+1/2} \mathbf{E}^{n+1/2} + (n\mathbf{u})^{n+1/2} \times \mathbf{B}^{n+1/2} \right]$$

- See ISSS-series books by Winske et al.: (1991, 2003, 2022/2023?)

Modern predictor²-corrector scheme



(Kunz et al., JCP 259, 154 (2014))

arXiv:2204.01676

Time-stepping: Whistler wave dispersion

- Inclusion of Hall-term gives extremely stiff waves:

$$\frac{\partial \mathbf{B}}{\partial t} = -\nabla \times \left(\frac{d_i}{n} (\mathbf{j} \times \mathbf{B}) \right) \longrightarrow \omega = id_i v_A k^2$$

Quadratic dispersion

- This sets stiff CFL:

$$k_{\max} \sim \frac{1}{\Delta x} \longrightarrow \Delta t_{\text{CFL}} \sim n (\Delta x)^2 / B$$

- Usually shows up first as numerical instabilities in near vacuum regions.
- Potential solutions:
 1. Sub-cycling,
 2. Electron inertia,
 3. Re-introduce speed of light,
 4. Implicit time-stepping.

Implicit time-stepping in hybrid-PIC methods

- Early efforts: implicit field solve (e.g. Hewett, 1980).
- More recently, fully implicit methods have been developed:
 1. Implicit scheme for electrostatic δF model (Sturdevant et al., J. Comp. Phys. 2016)
 2. Implicit scheme for electromagnetic full-F (Stanier et al., J. Comp. Phys. 2019).
 - Can take steps much larger than whistler Δt_{CFL} .
 - Exact conservation of momentum & energy.
 - Build on recent breakthroughs in implicit PIC methods (Markidis et al. J.Comp. Phys. 2011; Chen & Chacon J. Comp. Phys. 2011).

Numerical stability and dissipation

- Although space plasmas are close to collisionless, nonlinear numerical simulations typically need some dissipation for stability. Either via:
 1. Explicit terms in equations (“physical dissipation”).
 2. Upwinding of advective terms (implicit dissipation via discretization).
- Hybrid models usually follow 1) by adding **dissipation in Ohm’s law**:

$$\mathbf{E} = \mathbf{E}^* + \eta \mathbf{j} - \eta_H \nabla^2 \mathbf{j}$$

Frictionless E resistivity “Hyper-resistivity”

Hyper-resistivity: Why and what?

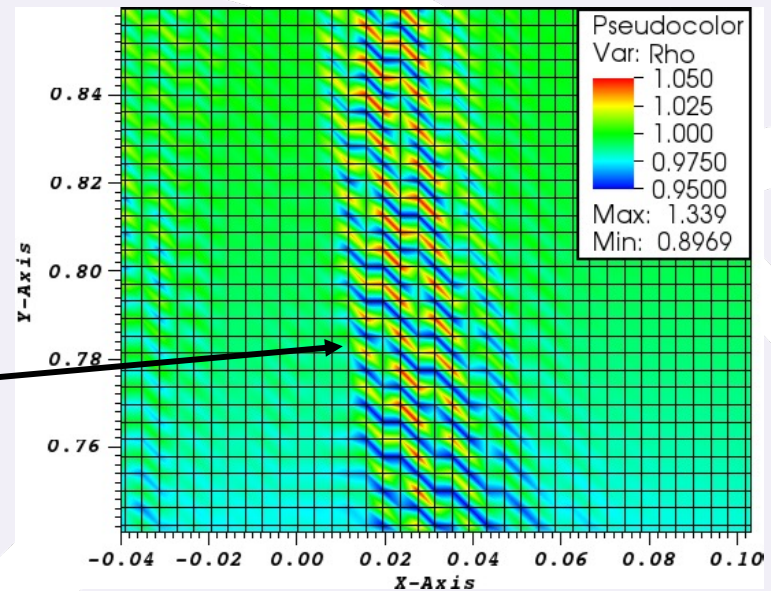
- Why? Hall-term is badly behaved!
- Consider damped Whistler dispersion:

$$\omega = -ik^2\eta - ik^4\eta_H \pm d_i v_A k_{\parallel} k.$$

- a) Balance 1 & 3: Both $\sim k^2$ - Can't set a dissipation scale – Whistler noise!
- b) Balance 2 & 3:

$$\lambda = \frac{2\pi}{k} = 2\pi \sqrt{\frac{\eta_H}{d_i v_A}}$$

➤ Want to set $\lambda \sim \Delta x$ to avoid whistler noise.



See e.g. Stanier, A., PhD Thesis (2013)

- What?
 - Similar form of an electron collisional viscosity: $\approx -\eta_H \nabla^2 \mathbf{u}_e$
 - But coefficient too large for space! Sometimes argued as “anomalous viscosity”.

Conservation when including frictional terms

1. Momentum conservation:

- Collisionless Vlasov: $\partial_t f_i + \nabla \cdot (f_i \mathbf{v}) + (q_i/m_i) (\mathbf{E}^* + \mathbf{v} \times \mathbf{B}) \cdot \nabla_v f_i = 0.$
- Collisional Ohms: $\mathbf{E} = \mathbf{E}^* + \eta \mathbf{j} - \eta_H \nabla^2 \mathbf{j}$
 - Total momentum conservation: Push with $\mathbf{E}^*.$

2. Energy conservation:

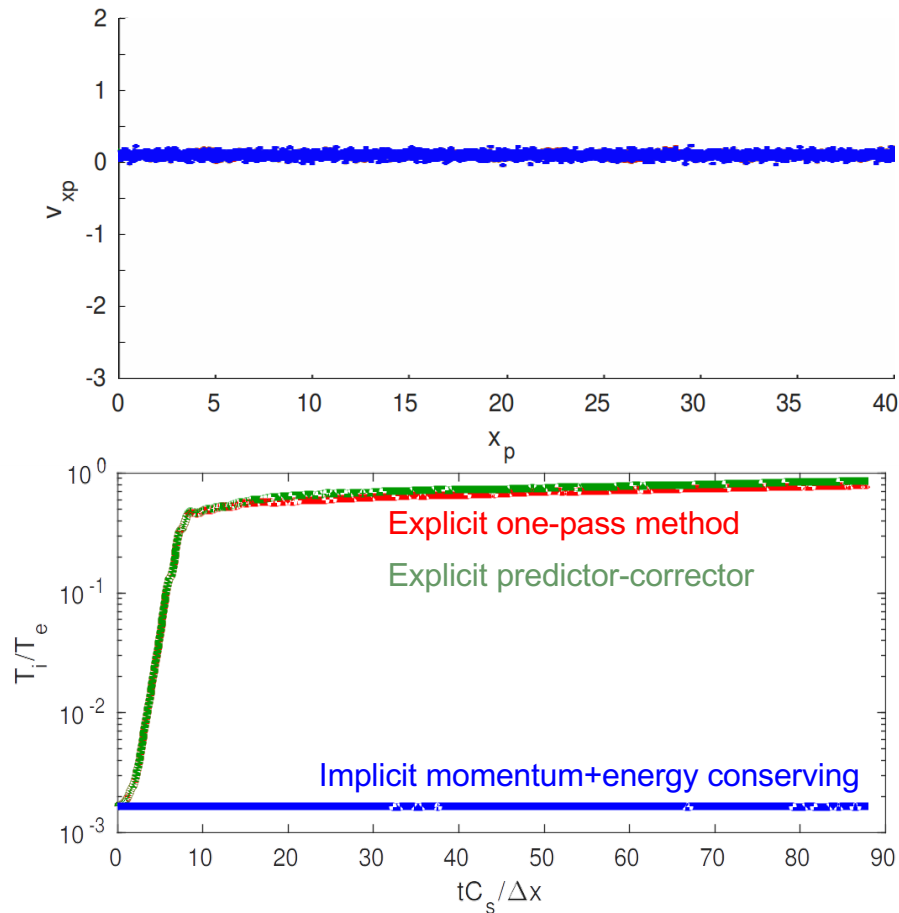
- Requires separate electron pressure equation with heating terms:

$$(\gamma - 1)^{-1} [\partial_t p_e + \nabla \cdot (\mathbf{u}_e p_e)] + p_e \nabla \cdot \mathbf{u}_e = H_e - \nabla \cdot \mathbf{q}_e.$$

$$\text{Frictional heating: } H_e = \eta j^2 + \eta_H \nabla j : \nabla j$$

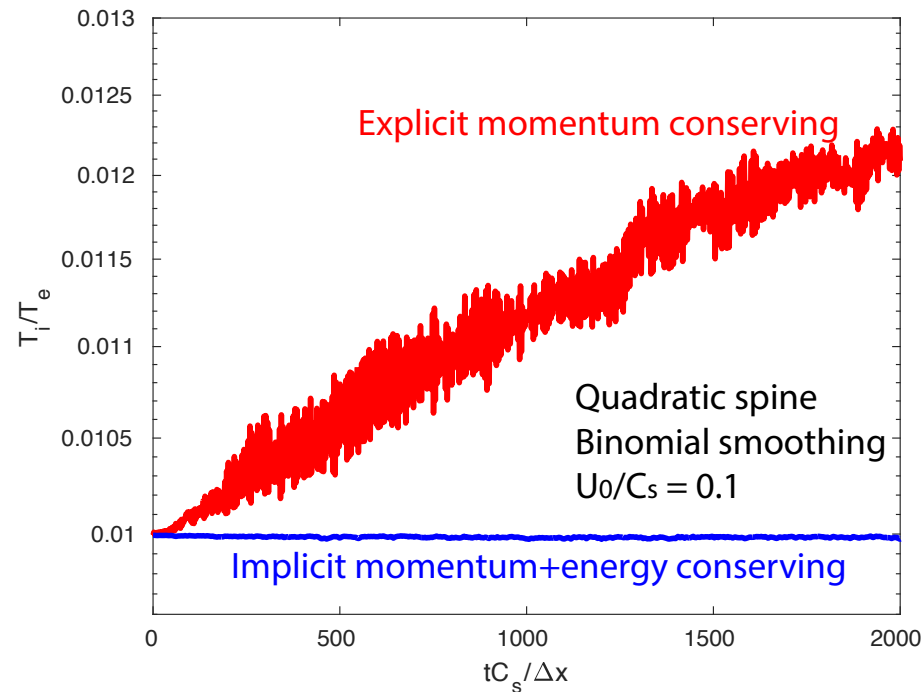
Finite grid instabilities for cold ions

Nearest Grid Point, no smooth, $U_0/C_s = 0.1$.



- Problem set-up: Cold ion beam moving through uniform spatial mesh.
- Non-conservative (explicit) schemes unstable for $T_i/T_e \ll 1$ regardless of spatial resolution (Rambo, J. Comput. Phys. 1995).
 - Precise threshold in T_i/T_e and beam velocity depends on shape-function.
 - NGP threshold \gg QS threshold.
- Cause unstable (exponential) heating of ions until some saturation value & also violates momentum conservation.
- Implicit momentum+energy conserving scheme stable w.r.t. these instabilities. (Stanier et al., J. Comp. Phys. 2019)

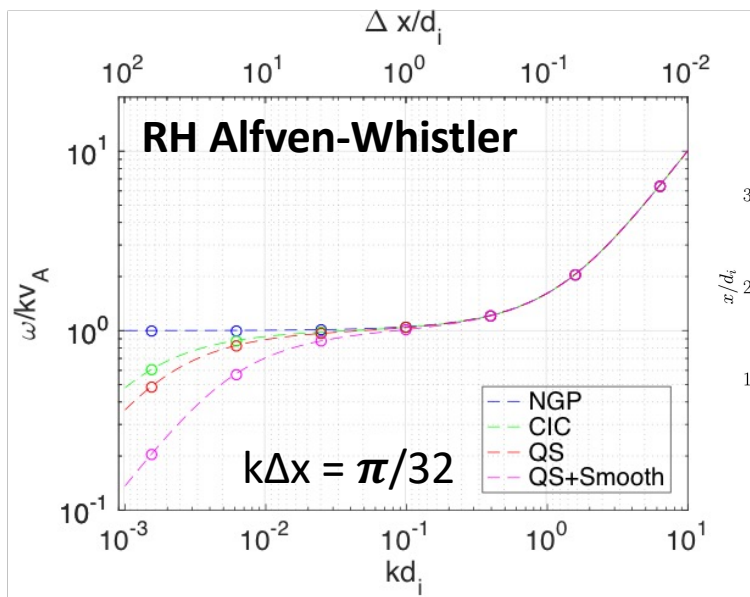
Stochastic heating when FGI stable



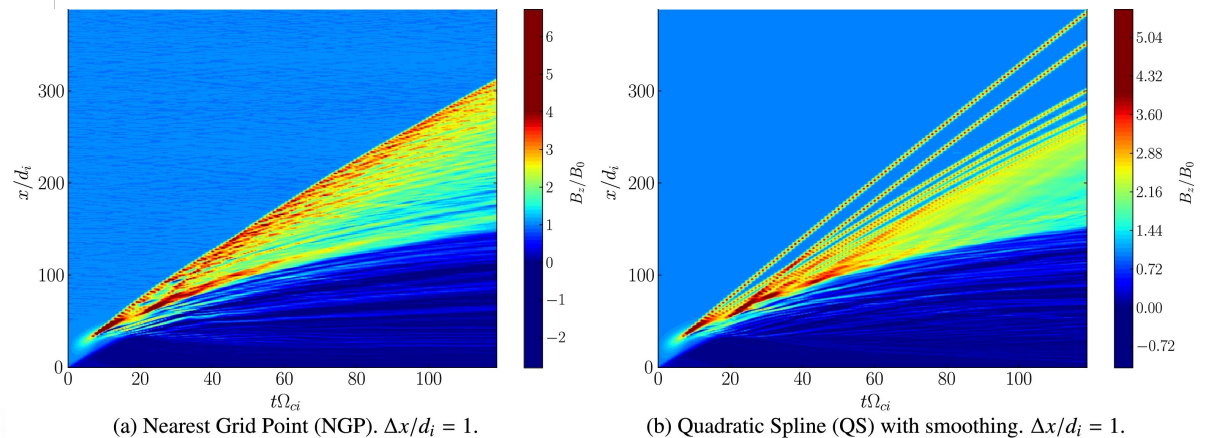
- For quadratic spline with smoothing, explicit becomes FGI stable for $T_i/T_e=0.01$.
- However, **significant stochastic heating for explicit** (Rambo, JCP 97).
- **Implicit scheme gives large improvement.**

Minimum mesh resolution requirements

- **“Hybrid cancellation problem”** due to different discretization of ions (particles) and electrons (fluid). *Stanier et al., JCP 420, 109705 (2020)*
 - The error depends on shape function & amount of smoothing.



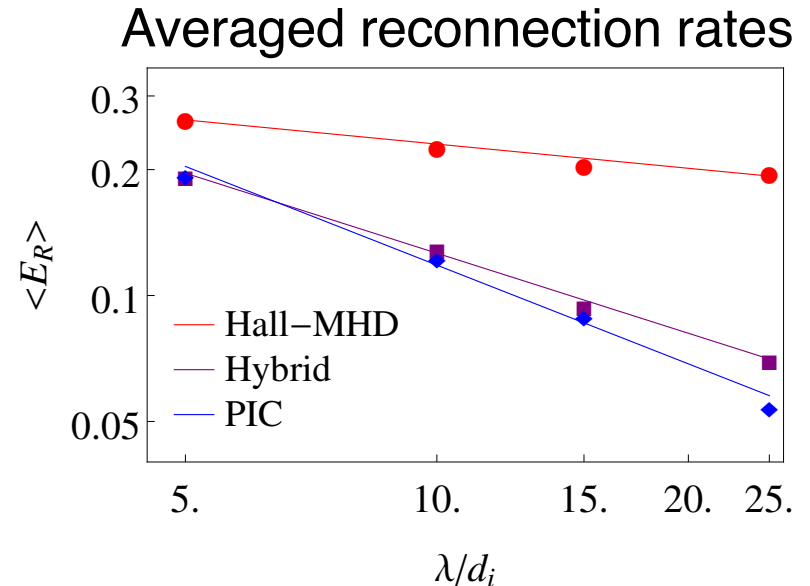
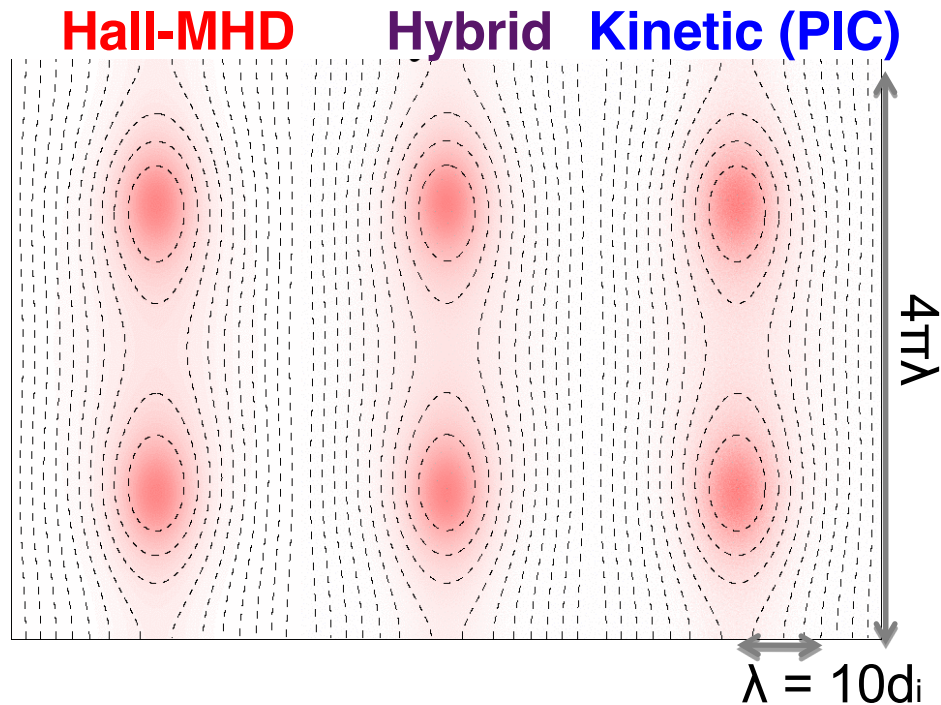
1D collisionless EM perpendicular shock generated by $M_A=5$ piston ions.



- NGP & no smoothing: can take $\Delta x \gg d_i$ (provided features resolved).
- In other cases: Need to resolve ion skin-depth ($\Delta x/d_i < 1$).

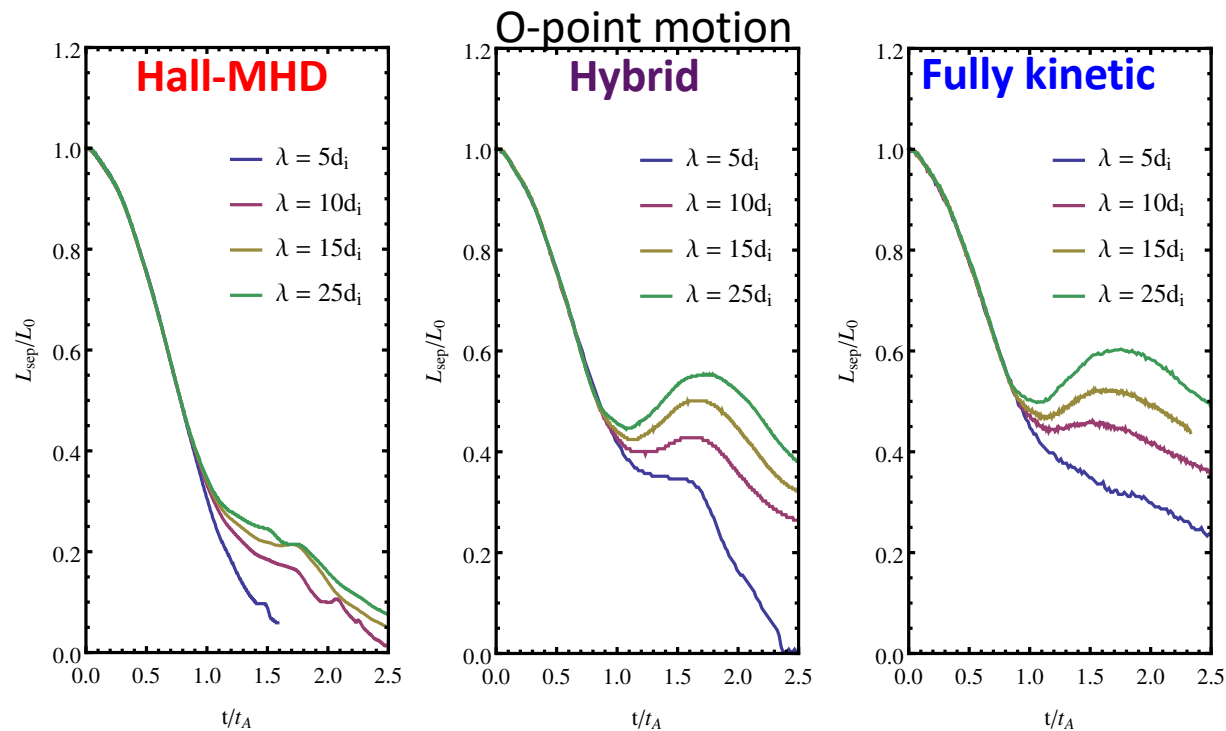
5. Applications of hybrid-PIC methods

Macro- to ion-scale coupling is important in reconnection



- **Hybrid** model minimum sufficient model to reproduce **fully kinetic** reconnection rates and macro-to-ion scale coupling (Stanier et al., PRL 2015, Ng et al., PoP 2016).
- Key ion kinetic physics missing from **Hall-MHD** fluid model.

Kinetic ion codes have different global behavior



- **Hybrid & fully kinetic:** O-point reversal (sloshing) for $\lambda \geq 10 d_i$.
- **Hall-MHD:** No clear reversal for $\lambda \leq 25 d_i$.
- Missing kinetic ion physics gives different global evolution.

Comparison with Hall-MHD

- Now if we similarly take moments of ion-Vlasov:

- 0th Continuity: $\partial_t n + \nabla \cdot (n \mathbf{u}_i) = 0,$

- 1st (i+e) momentum:

$$\partial_t(mn\mathbf{u}_i) + \nabla \cdot \left[mn\mathbf{u}_i\mathbf{u}_i - \mathbf{B}\mathbf{B}/\mu_0 + \mathbb{I}(B^2/2\mu_0) + \overline{\overline{\mathbf{P}}} \right] = \mathbf{0}.$$

Hall-MHD (constant T_{i0}/T_{e0})

$$\overline{\overline{\mathbf{P}}} = p_e(1 + T_{i0}/T_{e0})\mathbb{I}$$

Hybrid (single ion species)

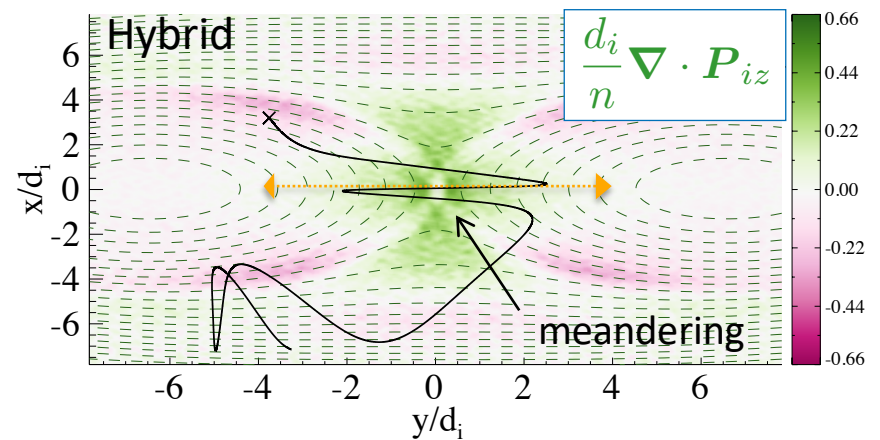
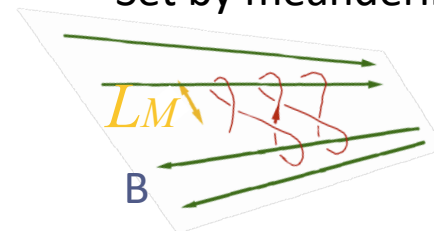
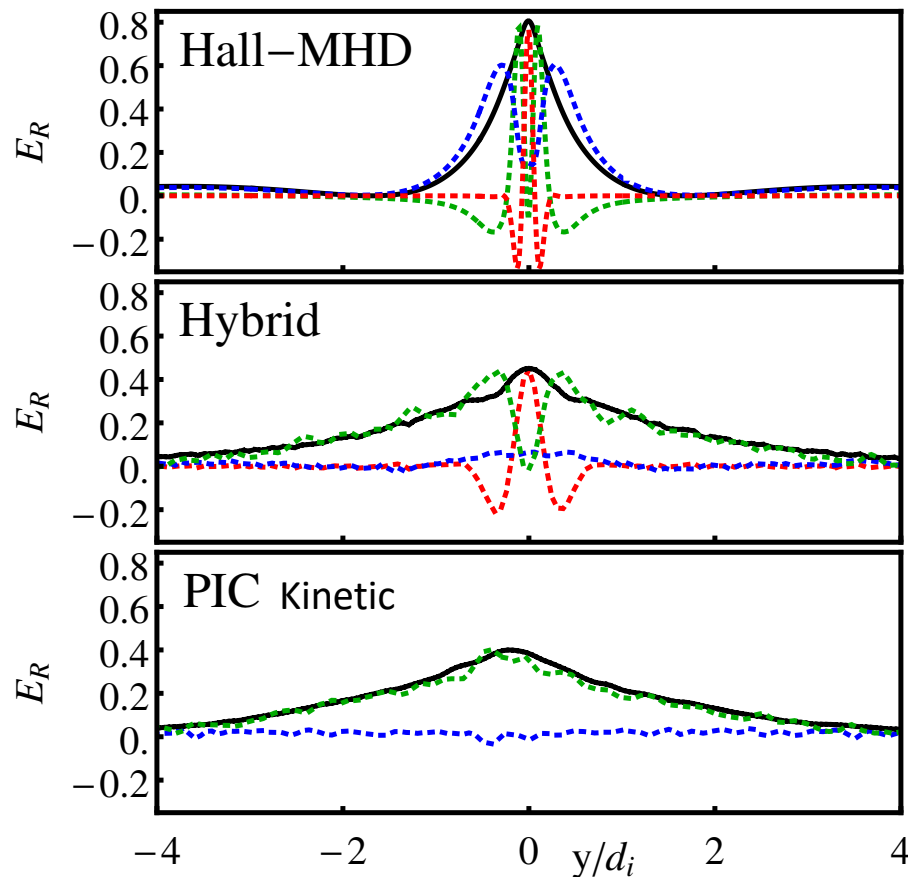
$$\overline{\overline{\mathbf{P}}} = p_e\mathbb{I} + \int m_i f_i \mathbf{w} \mathbf{w} d^3w$$

- Hall-MHD is a “cold-ion” model in the sense that it does not include ion finite Larmor radius (FLR) or other kinetic effects from warm distribution functions.

Missing physics: Finite ion-orbit effects

$$E'_z = (\mathbf{E} + \mathbf{v}_i \times \mathbf{B}) \cdot \hat{z} = \frac{d_i}{n} [\partial_t(nv_{iz}) + \nabla \cdot (nv_i v_{iz})] + \frac{d_i}{n} \nabla \cdot \mathbf{P}_{iz} + F_{\text{coll},z}$$

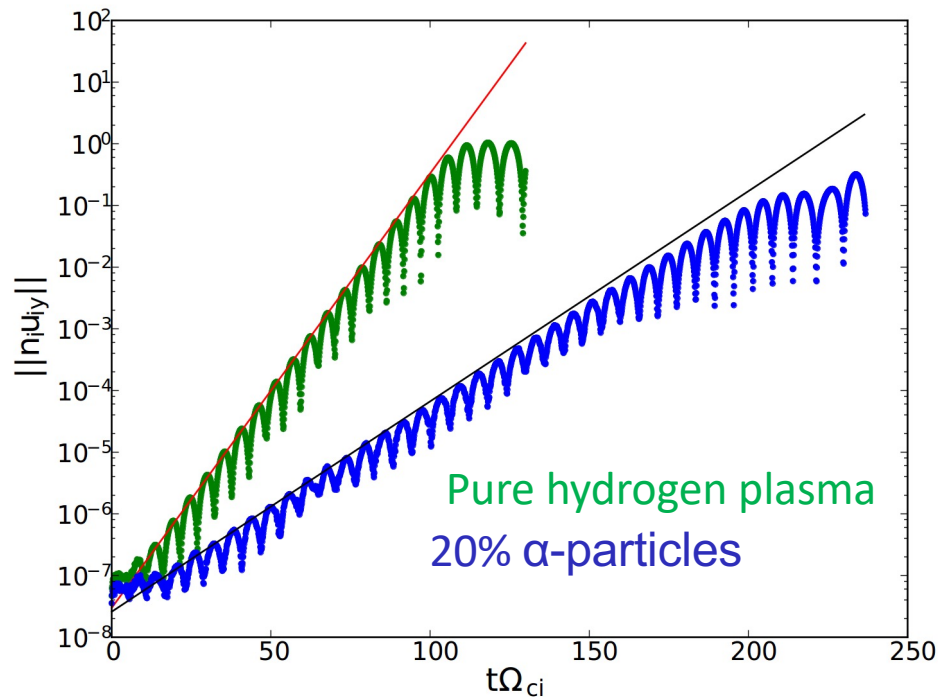
- Hybrid & fully kinetic:
 - Broader ion diffusion region.
 - Supported by non-gyrotropic \mathbf{P}_i .
 - Set by meandering-type orbits (Speiser 1965).



Proton Cyclotron Anisotropy Instability

■ Electromagnetic & multi-ion verification test:

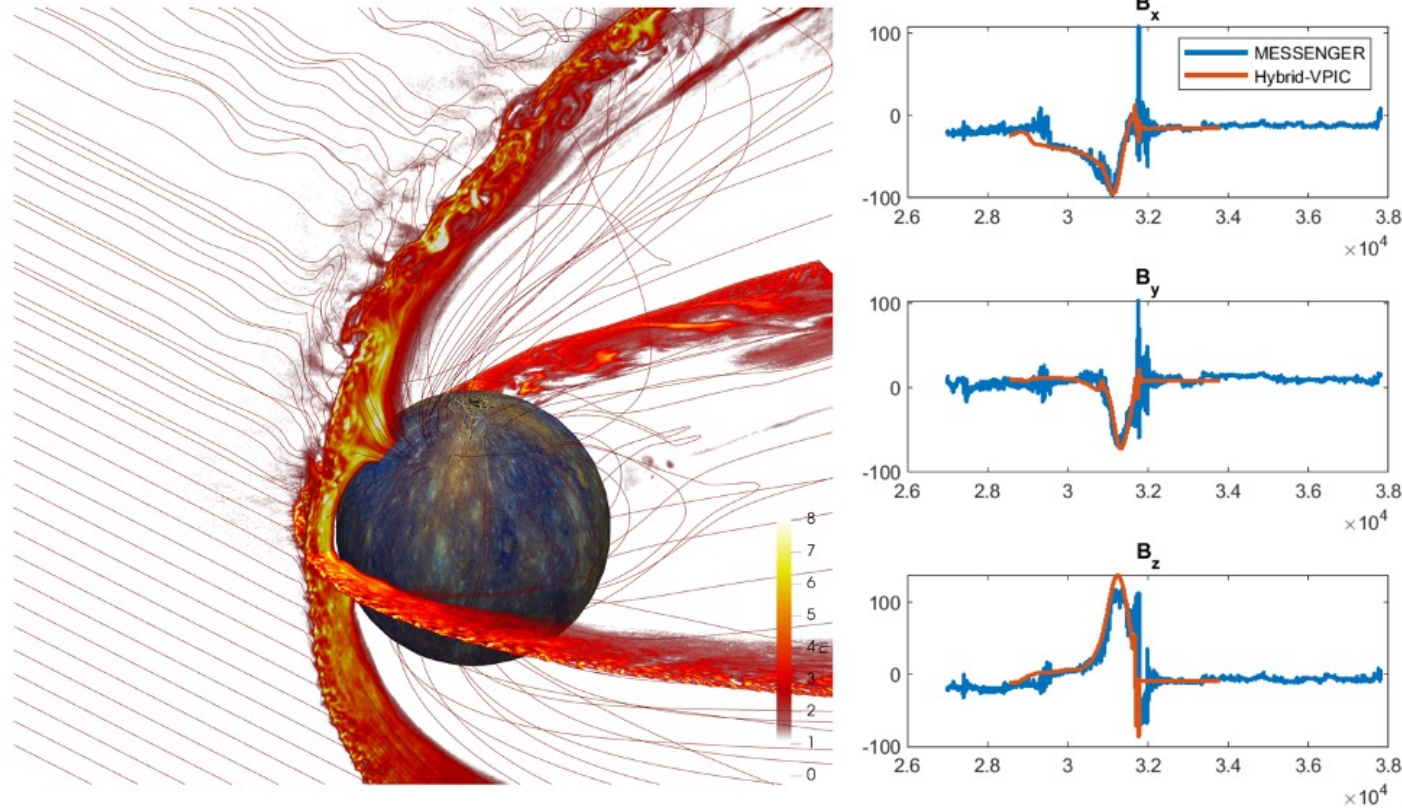
- 1D-3V electromagnetic instability driven by $p_{i\perp}/p_{i\parallel} > 1$.
- Maximum growth at $\mathbf{k}\times\mathbf{B}=\mathbf{0}$, finite real frequency.
- $k_x\Delta x = 0.065$, 50000 particles/cell, 2x binomial smooth, quiet start.



$$p_{\alpha\perp}/p_{\alpha\parallel} = p_{p\perp}/p_{p\parallel} = 3, \beta_{p\parallel} = 1, T_{\alpha\parallel}/T_{p\parallel} = 2, k_{\parallel}d_p = 0.6$$

Mercury

Results from Ari Le and Chuanfei Dong



- 3D global hybrid simulations of Mercury (using hybridVPIC).
- Comparison with MESSENGER M2 flyby.
- Formation of ion foreshock in quasi-parallel region.

Earth-scale simulations (2D&3D)

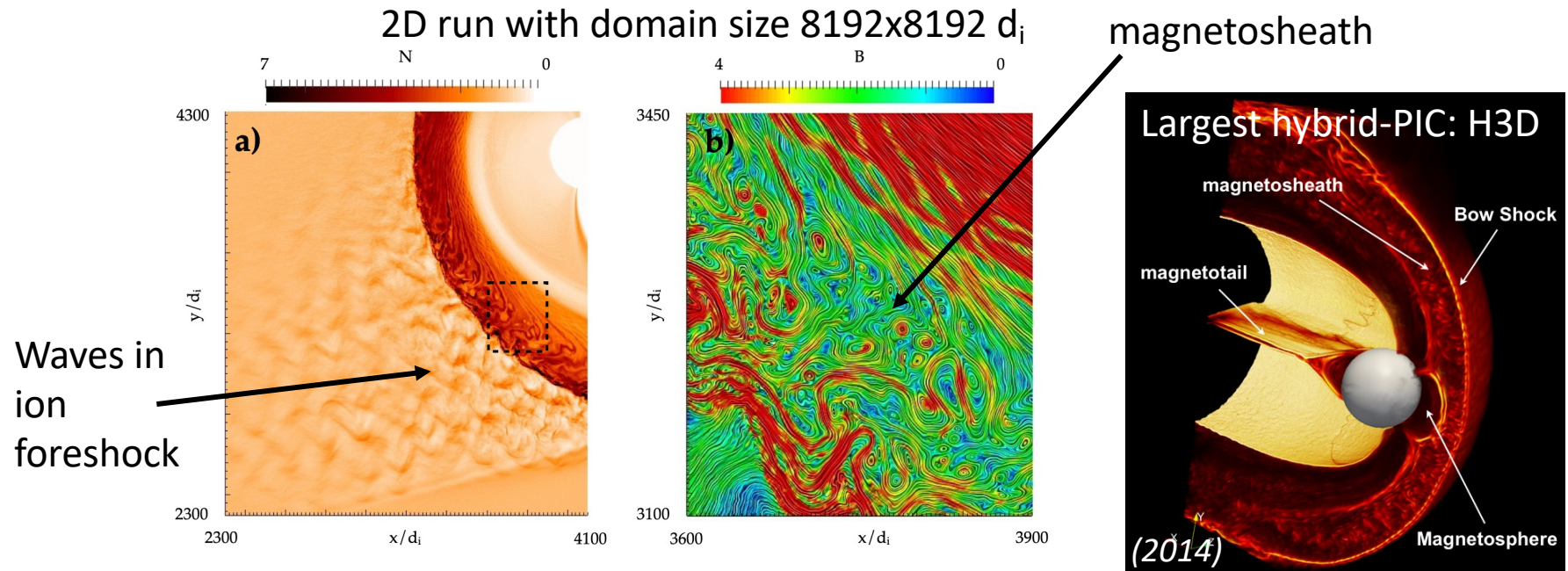


Fig. 4. Foreshock and magnetosheath turbulence in a 2D hybrid simulation describing interaction of solar wind injected from the left boundary with a dipolar field. Left: ion density; Right: LIC visualization of the magnetic field structure in a sub-region marked on the left panel. The scales are normalized to the ion inertial length d_i .

- (Karimabadi et al., Phys. Plasmas 2014): “The links between shocks, reconnection and turbulence”.

6. Fast particle kinetic hybrid model

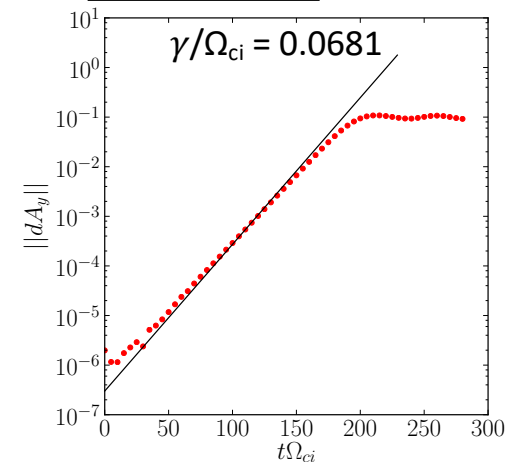
Fast-particle hybrid models

- Have greater use in magnetic fusion and astrophysics (cosmic rays) applications.
- Some recent interest to develop new models to solar particle acceleration (Drake et al. Phys. Plasmas 2019).
- MHD + fast particle using current-coupling:

$$\left\{ \begin{array}{l}
 \partial_t n + \nabla \cdot (n\mathbf{U}) = 0, \\
 n\partial_t \mathbf{U} + n\mathbf{U} \cdot \nabla \mathbf{U} + \nabla p = q_h n_h \mathbf{E}^* + \left(\frac{\nabla \times \mathbf{B}}{\mu_0} - q_h n_h \mathbf{U}_h \right) \times \mathbf{B}, \\
 \partial_t p + \nabla \cdot (p\mathbf{U}) + (\gamma - 1)p \nabla \cdot \mathbf{U} = 0, \\
 \mathbf{E}^* = \mathbf{E} - \eta \mathbf{j} = -\mathbf{U} \times \mathbf{B} \\
 \partial_t \mathbf{B} = -\nabla \times \mathbf{E}, \\
 \dots
 \end{array} \right.$$

$$\left\{ \begin{array}{l}
 \partial_t f_h + \mathbf{v} \cdot \nabla f_h + \frac{q_h}{m_h} (\mathbf{E}^* + \mathbf{v} \times \mathbf{B}) \cdot \nabla_v f_h = 0. \\
 n_h = \int f_h d^3v, \quad n_h \mathbf{u}_h = \int f_h \mathbf{v} d^3v.
 \end{array} \right.$$

Alfven wave destabilized by tenuous ion beam



$$\begin{aligned}
 &B = 1, \quad q_{\text{beam}} = m_{\text{beam}} = 1, \\
 &n_{\text{beam}}/n = 0.05, \quad v_{\text{th,beam}} = v_A, \\
 &v_{0,\text{beam}} = 2.5 v_A
 \end{aligned}$$

6. Hands-on session

Hybrid-VPIC

VPIC (Vector Particle-In-Cell) code originally developed by K. Bowers (Bowers et al., Physics of Plasmas, 2008) has been used for some of the largest 3D PIC simulations performed:

- E.g. 10 trillion particles and 10 billion cells.
- VPIC philosophy: Simple algorithms that scale and run extremely fast.

“Hybrid-VPIC”: New version implementing the quasi-neutral hybrid model:

- “Single pass” algorithm of H3D: Leapfrog particles, RK4 field-solve.

<https://doi.org/10.1063/5.0146529>

- Primary developer: Ari Le (Los Alamos National Laboratory).
- Open source:
<https://github.com/lanl/vpic-kokkos/tree/hybridVPIC>
- Vlasov-Maxwell PIC: <https://github.com/lanl/vpic>
<https://github.com/lanl/vpic-kokkos>

Hybrid-VPIC best practices

- For each simulation create a new folder in “scratch” (/ptmp/mpXX/folder) to run from.
- Build input deck against source every time changes are made “make”.
- Beware of dumping particle data: Can use massive amount of disk-space.
- In case of error, check log file e.g. “vpic.out”.
- Some things to try:
 - Decrease timestep.
 - Increase (or decrease) dissipation, e.g. hyper-resistivity.
- Contact: stanier@lanl.gov for questions/issues/bug reporting.
- Warning: Recently open sourced - check github for bug fixes! (and new features).

Proton Cyclotron Anisotropy Instability (PCAI)

- Electromagnetic instability driven by $P_{p\perp} > P_{p\parallel}$.
- Left hand polarization with resonant ions.
- Fastest growing mode has $kxB = 0$.
- Instability threshold:

$$\frac{P_{\perp}}{P_{\parallel}} - 1 \approx \frac{S}{\beta_{p\parallel}^{0.4}} \text{ with } S \sim 1.$$

- Collisionless wave-particle scattering reduces P_{\perp}/P_{\parallel} until saturation.

Gary, S. P. (1993). *Theory of space plasma microinstabilities* (No. 7). Cambridge university press.

PCAI input deck

1. Running and postprocessing data (see pcai/README).

- `cd pcai`
- `emacs -nw pcai.cxx`
- `make`
- `sbatch subslurm`
- `ftn -o translate_pcai translate_pcai.f90`
- `mkdir data`
- `./translate_pcai`
- `python ./plotsPCAI.py`

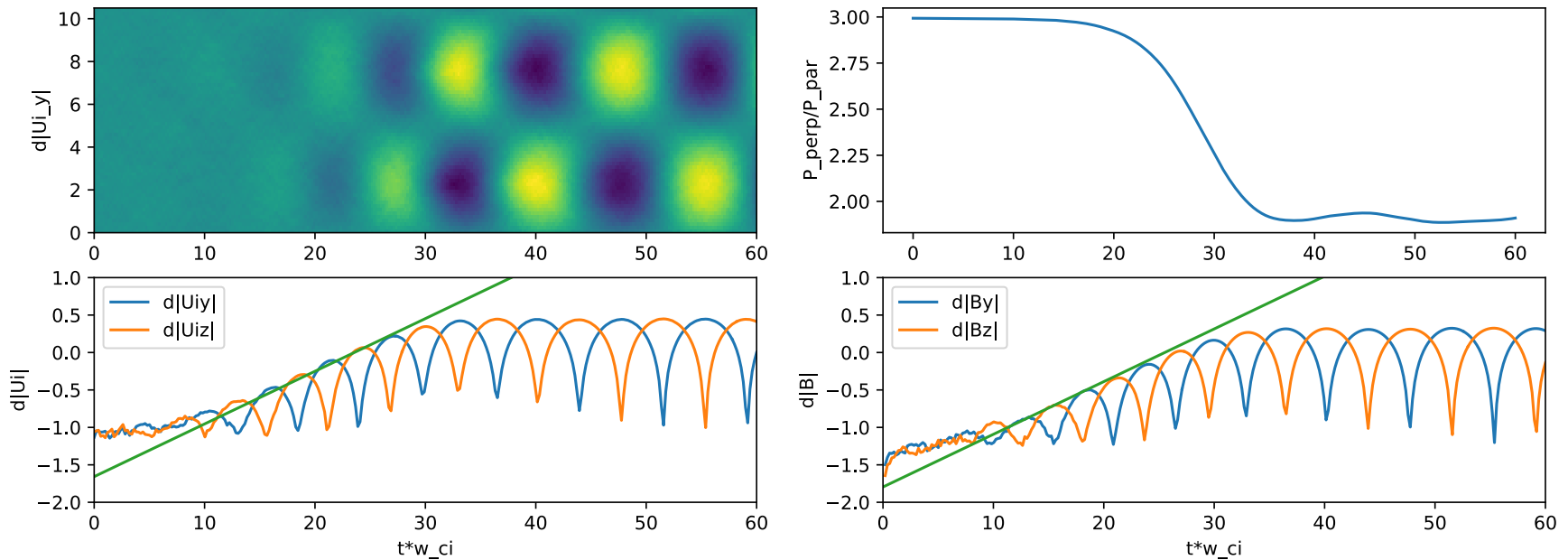
2. Nominal simulation parameters:

$$\beta_{p\parallel} = 1, \quad \frac{T_{p\perp}}{T_{p\parallel}} = 3, \quad \frac{L_x}{d_i} = 10.5, \quad \frac{T_e}{T_{p\parallel}} = 1, \quad \gamma = \frac{5}{3}.$$

3. Numerical parameters:

- 1D simulation, 64 cells, 10K particles/cell, $\Delta t \Omega_{ci} = 0.01$, dissipation=0.

PCAI results for nominal parameters



- Transverse velocity and magnetic field components grow from noise (LH Alfvén waves).
- Agree with linear theory for these parameters ($\gamma/\Omega_{ci}=0.162$).
- Pressure anisotropy decreases until saturation.

PCAI: Suggested exercises

- Compute growth rates across a range of beta & anisotropy.
 - Compare against linear solver. E.g. “HYDROS” by D. Told (New Journal Physics, 2016 - <https://github.com/dtold/HYDROS>).
- Advanced: Add a 20% density fraction of a minor species of alpha particles and find how this modifies the growth rate.
 - For parameters and results, see: Stanier, A., et al. (2019). A fully implicit, conservative, non-linear, electromagnetic hybrid particle-ion/fluid-electron algorithm. *Journal of Computational Physics*, 376, 597-616.

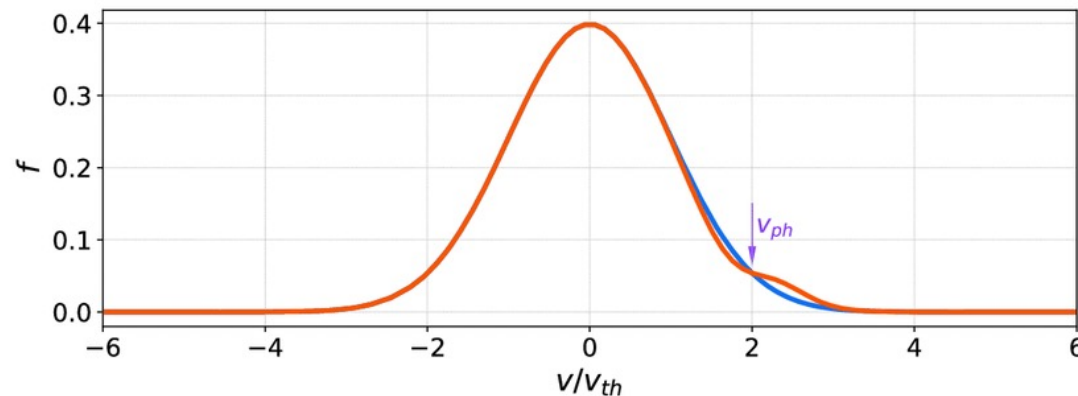
Landau Damped Ion Acoustic Wave

- Fundamental electrostatic mode in hybrid-PIC model: Ion Acoustic Wave.
 - Driven by pressure perturbation.
 - Fluid models (e.g. Hall-MHD): wave is undamped.
 - Hybrid-PIC: Landau resonance damps the wave & locally flattens ion VDF (analogous to electron LD of Langmir waves).

- Dispersion relation:

$$\frac{dZ(\zeta)}{d\zeta} = 2\tau, \quad \zeta \equiv (\omega - i\gamma) / kv_{th,i}$$

Plasma dispersion function
T_i/T_e



IAW input deck

1. Running and postprocessing data (see iaw/README).

- `cd iaw`
- `emacs -nw iaw.cxx`
- `make`
- `sbatch subslurm`
- `ftn -o translateIAW translateIAW.f90`
- `mkdir data`
- `srun -n 1 ./translateIAW`
- `python ./plotsIAW.py`

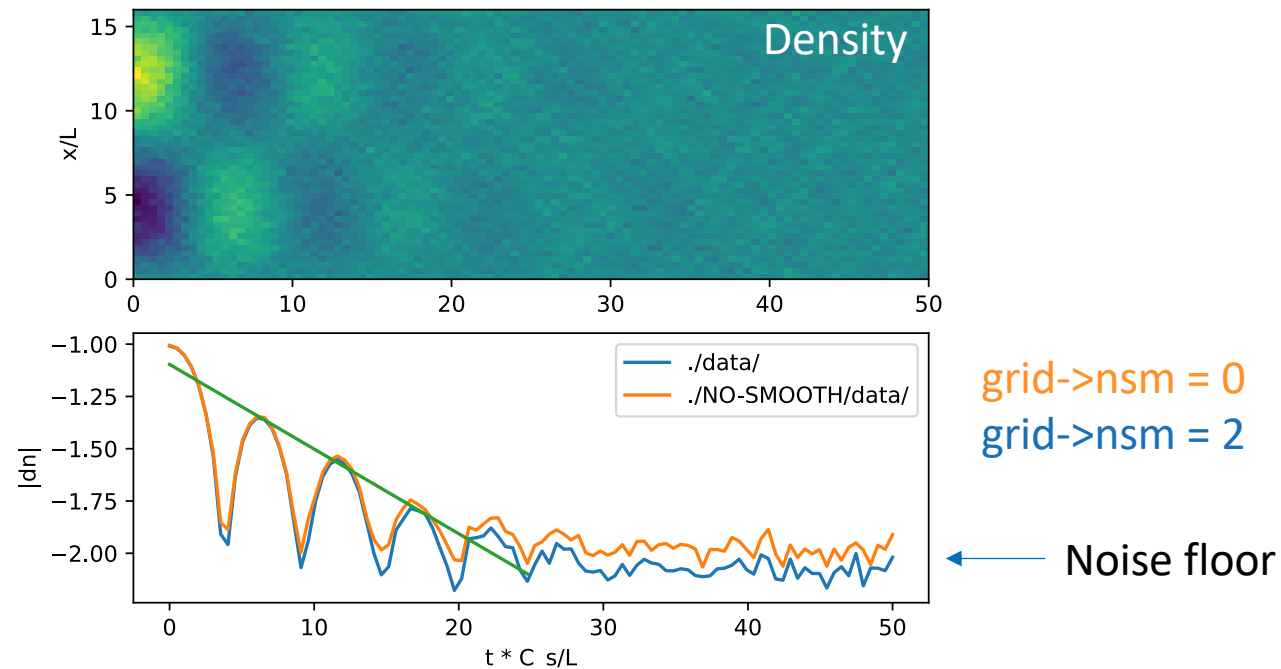
2. Nominal simulation parameters:

$$T_i = 1/3, \quad \gamma = 5/3, \quad c_s = \sqrt{\gamma T_e / m_i} = 1, \quad k_x = \pi/8, \quad \delta n = 2 \times 10^{-2}$$

3. Numerical parameters:

- 1D simulation, 48 cells, **150K particles/cell !**, $\Delta t = 0.02$, dissipation = 0.

Landau-damped IAW results for nominal parameters



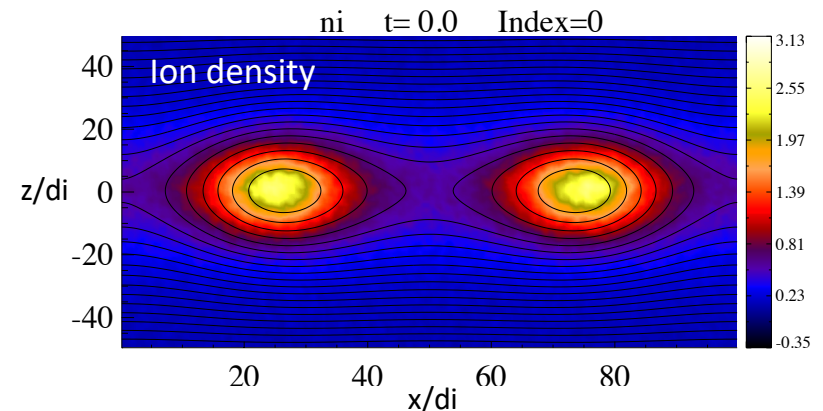
- Damping rate: $\gamma = -0.0932$.
- Initial perturbation damps to noise floor. Noise can be reduced by:
 1. Use more particles/cell (noise $\sim 1/\sqrt{N_p}$).
 2. Binomial smoothing/higher order shape functions (see figure).
 3. Using low-discrepancy quasi-Random numbers to seed particles (noise $\sim 1/N_p$).
 4. Most efficient: Delta-F

Landau-damped IAW: Suggested further exercises

- See how temperature ratio T_i/T_e influences damping rate.
- What happens when we set $T_i \rightarrow 0$? (Finite grid unstable).
 - How is this the numerical instability threshold influenced by binomial smoothing?
- What happens when we use a larger perturbation? E.g. $dn = 0.5$?:
 - Non-linear Landau damping (phase space vortex structure).

Magnetic reconnection island coalescence

- Magnetic islands 2D versions of flux-ropes.
- Self-driven reconnection problem:
 - Coupling of ideal island motion to micro-scale reconnection physics.
 - **Ion kinetic effects are crucial.**
- Unstable Fadeev island equilibrium:



- Magnetic field: $\mathbf{B} = \nabla \times \mathbf{A}$

$$A_y = -\lambda B_0 \ln [\cosh (z/\lambda) + \epsilon \cos (x/\lambda)],$$

- Density:

$$n = n_0(1 - \epsilon^2) / [\cosh (z/\lambda) + \epsilon \cos (x/\lambda)]^2 + n_b$$

- Pressure balance: $\beta = \frac{2\mu_0 n_0 k_B (T_{i0} + T_{e0})}{B_0^2} = 1$

Island coalescence: input deck

1. Running and postprocessing data.

- cd islands
- emacs -nw islands.cxx
- make
- sbatch subslurm
- ftn -o translate_islands translate_islands.f90
- mkdir data
- srun -n 1 ./translate_islands
- ftn -o ayprog ay_gda_integrate.f90
- ./ayprog
- **IDL gui:** “module load idl” “idl” “diagnostic”. **Python:** “mkdir figs” “python plotfigs.py”.
- “python ./rate.py”, “python opoint.py”.

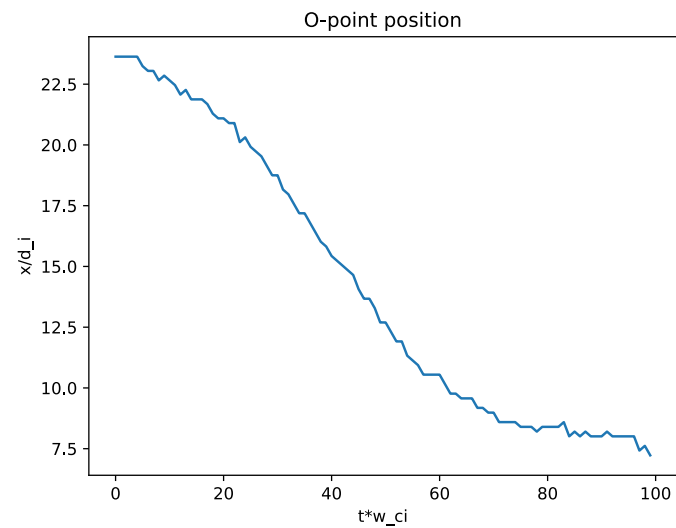
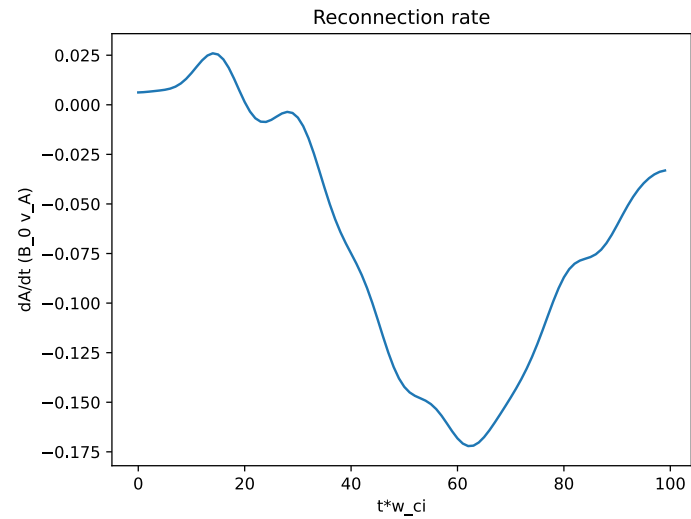
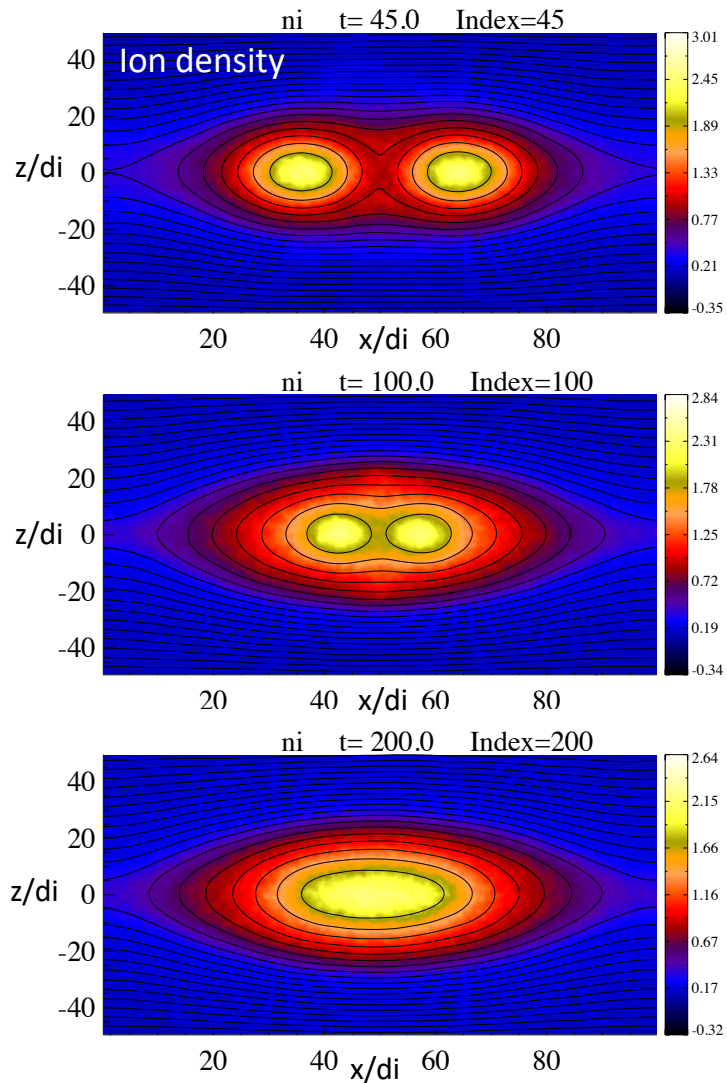
2. Nominal simulation parameters:

$$\lambda = 5d_i, \quad \epsilon = 0.4, \quad n_b = 0.2n_0, \quad T_i/T_e = 1, \quad \eta = 10^{-3}, \quad \eta_H = 5 \times 10^{-3}, \quad \gamma = 1$$

3. Numerical parameters:

- 2D simulation, 256x128 cells, 50 particles/cell, $\Delta t * \Omega_{ci} = 0.005$.

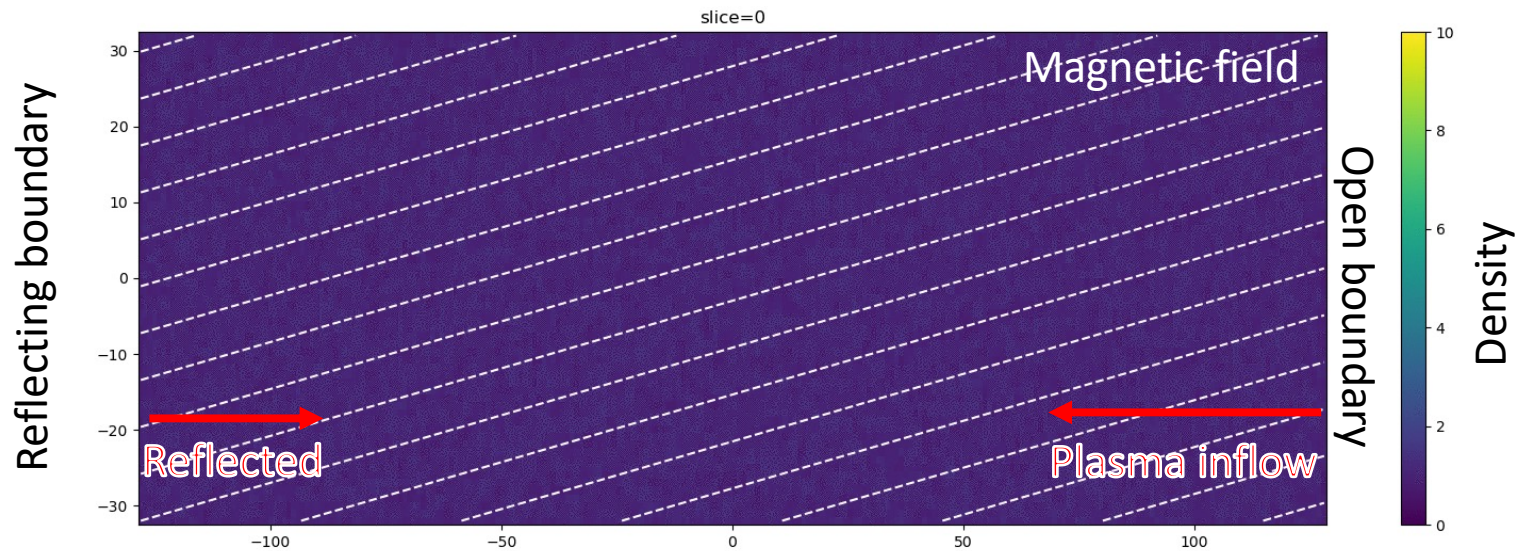
Island coalescence problem: Results



Island coalescence problem: Suggested exercises

- For these: Increase: resolution & particles/cell. Decrease: dissipation & timestep.
 1. Vary ratio of T_i/T_e and see how this modifies the reconnection rate and O-point motion (see Stanier et al., "Role of ion kinetic physics in the interaction of magnetic flux ropes", Phys. Rev. Lett. (2015)).
 2. How does the addition of a "guide" field change the structure of the diffusion region and the reconnection rate? (see Stanier et al., "The role of guide-field in magnetic reconnection driven by island coalescence", Phys. Plasmas (2017)).
 3. How do the maximum and average (merge time) rates of magnetic reconnection scale with the size of the islands (increasing λ/d_i).

Collisionless shock



2D magnetospheric shock problem:

- $M_A = 11.4$ injected from right (open) boundary.
- Reflects from left boundary to drive collisionless shock.

Thank you for your attention!

Advertisement: LANL Space Weather Summer School

- Flyer from 2023.
- **Next summer school in 2025**
– look out for
announcements.
- Deadline for applications
usually in January.



The flyer features a dark background with a vibrant, colorful visualization of a magnetosphere or solar wind interaction on the right side. The text is arranged in a structured layout with clear headings and bullet points. At the bottom, there are logos for the Los Alamos Space Weather Summer School, the Los Alamos National Laboratory, and the Center for Space and Earth Sciences (CSES), along with a QR code.

Los Alamos Space Weather Summer School June 5 - July 28, 2023

Los Alamos National Laboratory, Los Alamos New Mexico
<http://swx-school.lanl.gov>

The Los Alamos Space Weather Summer School is accepting applications for its 2023 session. Sponsored by the Center for Space and Earth Sciences at Los Alamos National Laboratory (LANL), this summer school brings together top space science graduate students and LANL space scientists to work on challenging space weather research. Students receive a prestigious Vela Fellowship (worth \$13,000 to cover travel and living expenses), technical training, and opportunities for professional development.

Lectures

The lectures will encompass three main themes. The first part will be an overview of basic space physics concepts geared towards understanding how the magnetosphere works and how it is driven. This will include the use of modeling tools to explore the same concepts in a more quantitative way, exposing the strengths and weaknesses of available models. The second part of the lectures will bring these concepts together to explore how new space missions could be devised to help resolve longstanding scientific questions. The third part of the lectures will highlight on-going space science related activities at LANL and will include a "career day" to convey job opportunities and desirable skill-sets for a career in space physics. Lectures will be coordinated with "labs" to get more hands-on experience. Space data analysis and modeling will be the main themes of the labs. Several field trips will be organized to visit Los Alamos facilities and historic sites (examples could include LANSCE, electron accelerators, visualization and high-performance computing labs, etc.).

Research projects

A unique aspect of the Los Alamos Space Weather Summer School is its emphasis on scientific research projects. Students team up with LANL mentors to work on unresolved scientific problems in space physics. LANL is engaged in a wide variety of space-physic activities and offers a host of exciting research projects. Check online at <http://swx-school.lanl.gov> for a list of current and past projects. Students can also propose their own ideas, which might include topics from their PhD thesis (contact the Space Weather Summer School management to find a suitable match to a LANL mentor). In the past, the majority of these projects led to presentations at major international conferences and, in some cases, to publications in peer-reviewed journals.

Students

Open to U.S. and foreign graduate students currently enrolled in PhD programs in space physics, planetary science, aerospace engineering, or related fields. Acceptance is based primarily on student's academic record, list of publications and presentations, letter of nomination, and content of cover letter. Preference will be given to students pursuing careers in the space sciences and who have completed at least their first year of graduate school, but students in any year may apply.

More information and how to apply:
Please visit the Summer School website at <http://swx-school.lanl.gov> for more information. Application materials should be sent to swx-school@lanl.gov.

Applications will be accepted starting early December 2022 and are due early January 2023. Acceptance notification will be sent out in early February 2023.

Questions ? Email us at swx-school@lanl.gov

Please include the following materials with your application:

1. Cover letter describing your research interests, why you would benefit from the summer school, potential project ideas (they can be related to your thesis work), and mentor requests
2. A current CV including full list of publications and presentations
3. Your undergraduate and graduate transcripts
4. A brief description of your PhD program and current progress
5. A nomination letter from your advisor
6. Two additional letters of reference if not enrolled in U.S. PhD program

Los Alamos Space Weather Summer School
Los Alamos NATIONAL LABORATORY
CSES
Los Alamos



Advertisement: Post-doc position in applied mathematics and plasma physics

Algorithm development:

- Low-noise particle-in-cell methods,
- low-rank methods,
- hybrid fluid-kinetic methods.

For applications of:

- Space weather,
- Magnetic confinement fusion,
- Inertial confinement fusion.



Apply at: <https://lanl.jobs/Search/JobDetails/postdoctoral-researcher-in-applied-mathematics--plasma-physics> (email stanier@lanl.gov for details).

LANL postdoc program:

<https://collaboration.lanl.gov/postdoctoral-research/>