

The weak-signal Bayes-factor approximation

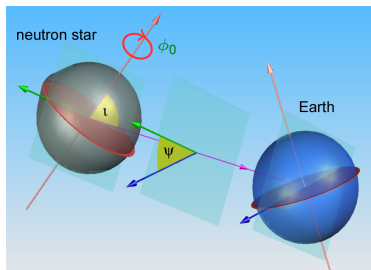
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Continuous gravitational waves and neutron stars @ Hannover
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Signal **amplitude** parameters



👉 **Amplitude** parameters:

$$\mathcal{A} = \{h_0, \underbrace{\cos \iota, \psi, \phi_0}_{\text{Euler angles}}\}$$

“radius” ($h_0 \geq 0$) + 3 Euler angles

\mathcal{A} generally unknown

Any spin orientation equally likely (isotropic):

$$P(\cos \iota | l) = \text{unif}(-1, 1)$$

$$P(\psi | l) = \text{unif}(0, 2\pi)$$

Any initial phase equally likely:

$$P(\phi_0 | l) = \text{unif}(0, 2\pi)$$

👉 “physical prior” $P(h_0, \cos \iota, \psi, \phi_0 | l) = \text{const} \times \underbrace{P(h_0 | l)}_?$

Part I: Pre-history

From $\mathcal{F}(x)$ - to $\mathcal{B}(x)$ -statistic

[Jaranowski, Królak, Schutz PRD58(1998)]

“cartesian” JKS coordinates $\mathcal{A}^\mu = \mathcal{A}^\mu(h_0, \cos \iota, \psi, \phi_0)$

$$\mathcal{A}^1 = 0.5h_0(1 + \cos^2 \iota) \cos 2\psi \cos \phi_0 - h_0 \cos \iota \sin 2\psi \sin \phi_0$$

$$\mathcal{A}^2 = 0.5h_0(1 + \cos^2 \iota) \sin 2\psi \cos \phi_0 + h_0 \cos \iota \cos 2\psi \sin \phi_0$$

$$\mathcal{A}^3 = -0.5h_0(1 + \cos^2 \iota) \cos 2\psi \sin \phi_0 - h_0 \cos \iota \sin 2\psi \cos \phi_0$$

$$\mathcal{A}^4 = -0.5h_0(1 + \cos^2 \iota) \sin 2\psi \sin \phi_0 + h_0 \cos \iota \cos 2\psi \cos \phi_0$$

 Signal strain @ detector

$$h(t; \mathcal{A}; \lambda) = \sum_{\mu=1}^4 \mathcal{A}^\mu h_\mu(t; \lambda)$$

$$h_1(t; \lambda) = a(t; \vec{n}_{\text{sky}}) \cos \phi(t; \lambda), \quad h_2(t; \lambda) = b(t; \vec{n}_{\text{sky}}) \cos \phi(t; \lambda)$$

$$h_3(t; \lambda) = a(t; \vec{n}_{\text{sky}}) \sin \phi(t; \lambda), \quad h_4(t; \lambda) = b(t; \vec{n}_{\text{sky}}) \sin \phi(t; \lambda)$$

Phase-evolution parameters: $\lambda = \{\vec{n}_{\text{sky}}, \text{freq}, \text{spindown}, \text{binary}, \dots\}$

(here: λ assumed known)

Likelihood ratio $\mathcal{L}(x; \mathcal{A}, \lambda)$

\mathcal{H}_G : data = Gaussian noise: $x(t) = n(t)$

\mathcal{H}_S : data = Gaussian noise + signal: $x(t) = n(t) + h(t; \mathcal{A}, \lambda)$

Goal: given $x(t)$ → “optimally” decide between \mathcal{H}_G and \mathcal{H}_S

→ likelihood ratio: Gaussian in \mathcal{A}^μ

$$\mathcal{L}(x; \mathcal{A}) \equiv \frac{P(x|\mathcal{H}_S, \mathcal{A})}{P(x|\mathcal{H}_G)} = \exp \left[-\frac{1}{2} \mathcal{A}^\mu \mathcal{M}_{\mu\nu} \mathcal{A}^\nu + \mathcal{A}^\mu x_\mu \right]$$

4 matched-filter outputs $x_\mu(\lambda) \equiv (x|h_\mu(\lambda)) \in \mathbb{R}$

4x4 antenna-pattern matrix:

$$\mathcal{M}_{\mu\nu}(\vec{n}_{\text{sky}}) \equiv (h_\mu|h_\nu) = \underbrace{(S_n^{-1} T)}_{\gamma \equiv \text{"quality x quantity"}} \begin{pmatrix} A & C & & \\ C & B & & \\ & & O & \\ & & & A & C \\ & & & C & B \end{pmatrix}$$

matched-filter scalar product: $(x|y) \equiv 2S_n^{-1} \int x(t) y(t) dt = 2\gamma \langle xy \rangle$

$A \equiv \langle a^2 \rangle$, $B \equiv \langle b^2 \rangle$, $C \equiv \langle ab \rangle$

1998: Maximum-likelihood: \mathcal{F} -statistic

☞ analytically maximize \mathcal{L} over \mathcal{A}^μ : [Jaranowski, Królak, Schutz PRD58(1998)]

$$\mathcal{F}(x) \equiv \max_{\{\mathcal{A}^\mu\}} \ln \mathcal{L}(x; \mathcal{A}^\mu) = \frac{1}{2} x_\mu \left(\mathcal{M}^{-1} \right)^{\mu\nu} x_\nu$$

$$2\mathcal{F}(x) = 2D^{-1} \left[AB(\mathcal{F}_A(x) + \mathcal{F}_B(x)) - 2C^2 \mathcal{F}_C(x) \right]$$

$$2\mathcal{F}_A(x) \equiv \frac{x_1^2 + x_3^2}{\gamma A}, \quad 2\mathcal{F}_B(x) \equiv \frac{x_2^2 + x_4^2}{\gamma B}, \quad 2\mathcal{F}_C(x) \equiv \frac{x_1 x_2 + x_3 x_4}{\gamma C}$$

$$\text{sub-determinant } D = \det \begin{pmatrix} A & C \\ C & B \end{pmatrix} = AB - C^2$$

☞ Long thought to be "optimal", until ...

Neyman-Pearson-Searle theorem [Searle arXiv:0804.1161]

☞ Bayes factor is the optimal statistic (if drawing from prior)

$$B_{S/G}(x) \equiv \frac{P(x|\mathcal{H}_S)}{P(x|\mathcal{H}_G)} = \int \mathcal{L}(x; \mathcal{A}) P(\mathcal{A}|I) d^4 \mathcal{A}$$

- “ \mathcal{F} -statistic prior” $P(\mathcal{A}^\mu) = \text{const}$ ☞ $B_{S/G} \propto e^{\mathcal{F}}$

(but prior is unphysical: $P(h_0, \cos \iota, \psi, \phi_0) \propto h_0^3 (1 - \cos^2 \iota)^3$)

- “physical prior” $P(h_0, \cos \iota, \psi, \phi_0) = \text{const} \times P(h_0)$

☞ can integrate ϕ_0 ✓

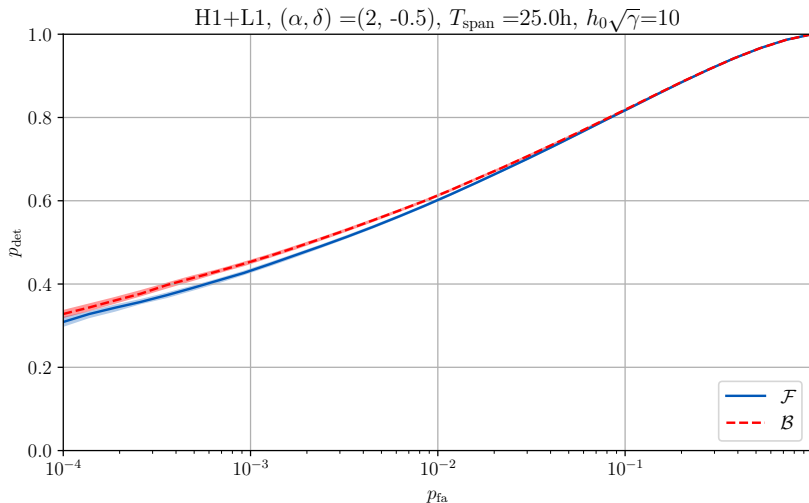
☞ if $P(h_0) = \text{const} \Rightarrow$ can integrate h_0 ! ✓

☞ “ \mathcal{B} -statistic”: 2D integral [Prix,Krishnan CQG26(2009)], [Whelan+ CQG31(2014)]

$$\mathcal{B}(x) \propto \int \frac{1}{\varrho} e^\Theta I_0(\Theta) d\cos \iota d\psi$$

$$\Theta(x; \cos \iota, \psi) \equiv \frac{q^2}{4\varrho^2}, \quad q = q(x; \cos \iota, \psi), \quad \varrho = \varrho(\cos \iota, \psi)$$

$\mathcal{B}(x)$ beats $\mathcal{F}(x)$



various attempts to solve/approximate $\mathcal{B}(x)$: [Dergachev PRD85(2012)][Whelan+ CQG31(2014)][Haris,Pai PRD96(2017)][Dhurandhar+ arXiv:1707.08163][Bero,Whelan CQG36(2019)][Wette Universe7(2021)]

Part II: Recent surprises

Short segments and segment weights

Surprise 1: dominant-response \mathcal{F}_{AB} -statistic [Covas, Prix PRD105(2022)]

Fully Coherent Search

Coherent + Coherent + Coherent + Coherent + Coherent

Sum = Semicoherent Search

All-sky binary search \Rightarrow very short segments $T_{\text{seg}} \sim \mathcal{O}(100\text{s})$

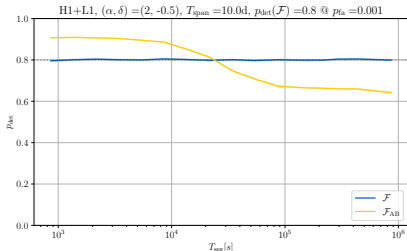
\Rightarrow degenerate $\mathcal{F}(x)$ in certain \vec{n}_{sky} where $D \rightarrow 0$

\Rightarrow maximum-likelihood $\mathcal{A}_{\text{ML}}^{\mu} = \mathcal{M}^{-1\mu\nu} x_{\mu}$ fails!

Fix: “dominant-response” statistic:

$$\mathcal{F}_{AB}(x) \equiv \begin{cases} \mathcal{F}_A(x) & \text{if } A > B \\ \mathcal{F}_B(x) & \text{otherwise} \end{cases}$$

\Rightarrow beats $\mathcal{F}(x)$ for short segments by up to $\sim 19\%!!$

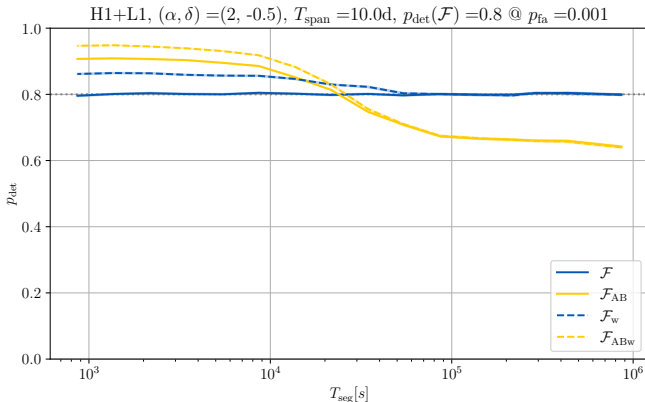


Surprise 2: segment weighting [Covas, Prix PRD106(2022)]

Can improve statistics by (empirical) *per-segment weights* w_ℓ :

("data factor" $\gamma = S_n^{-1}T$)

- $\hat{\mathcal{F}}_w \propto \sum_{\ell=1}^{N_{\text{seg}}} (\gamma (A + B))_\ell \mathcal{F}_\ell$
- $\hat{\mathcal{F}}_{ABw} \propto \sum_{\ell=1}^{N_{\text{seg}}} \left(\gamma \left(Q + \frac{C^2}{Q} \right) \right)_\ell \mathcal{F}_{AB,\ell}$ (with $Q = \max[A, B]$)

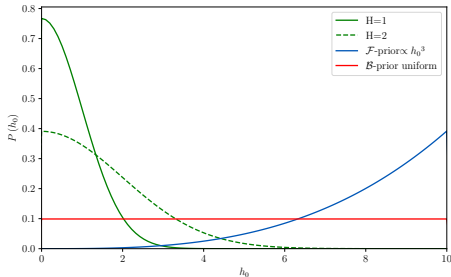


Part III Bayesian

Is there a ~~better~~ way?

Half-normal distribution with h_0 -“scale” H

$$P(h_0|I, H) = \frac{\sqrt{2}}{\sqrt{\pi} H} e^{-\frac{h_0^2}{2H^2}} \quad \text{for } h_0 \geq 0$$



- ✓ favors weak signals over strong
- ✓ proper prior
- ✓ $H \rightarrow \infty$ ~~→~~ reduces to uniform- h_0
- ✓ conjugate to \mathcal{L} : can integrate $h_0!$

“relative scale” $H_* \equiv H\sqrt{\gamma}$

$$\mathcal{B}(x; H_*) = \frac{1}{\pi} \int d\cos\iota d\psi \frac{1}{\sqrt{1 + H_*^2 \varrho^2}} e^{\Theta} I_0(\Theta)$$

$$\Theta(x; \cos\iota, \psi) \equiv \frac{H_*^2 \varrho^2(x; \cos\iota, \psi)}{4(1 + H_*^2 \varrho^2(\cos\iota, \psi))}$$

Two interesting limits:

- strong-signal limit: $H_* \gg 1$ → recover const- h_0 $\mathcal{B}(x)$ -stat
- weak-signal limit: $H_* \ll 1$: Taylor-expand in H_* !

(coherent sensitivity estimate $h_0 \Big|_{\substack{\rho_{\text{det}}=90\% \\ \rho_{\text{fa}}=1\%}} \sim 11.4/\sqrt{\gamma}$ → $h_0^{90\%} \sqrt{\gamma} \sim \mathcal{O}(10) \ll 1$)

Taylor-expand $\mathcal{B}(x; H_*)$ in $\mathcal{O}(H_*)$:

$$\mathcal{B}(x) = 1 + \frac{H_*^2}{4} \langle q^2 - 2\varrho^2 \rangle_{\cos \iota, \psi} + \mathcal{O}(H_*^4)$$

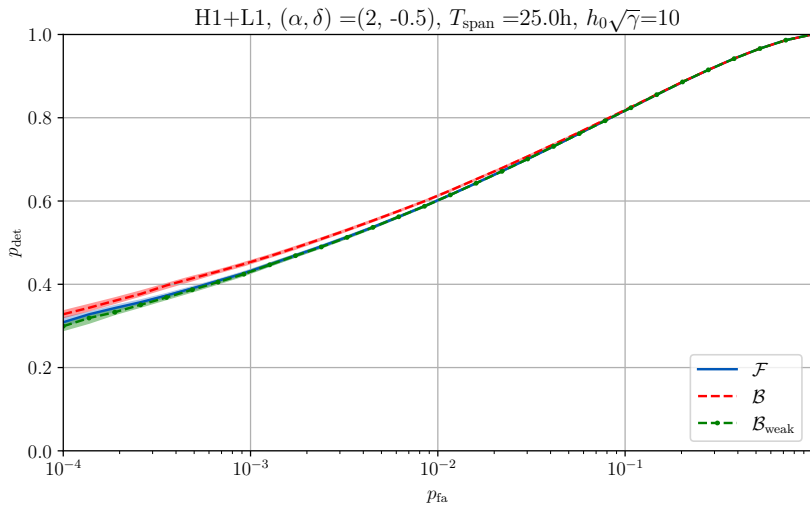
👉 Weak-signal limit Bayes factor $\mathcal{B}_{\text{weak}}(x)$

$$\ln \mathcal{B}_{\text{weak}}(x) \equiv \frac{H^2}{10} \gamma [A(2\mathcal{F}_A(x) - 2) + B(2\mathcal{F}_B(x) - 2)]$$

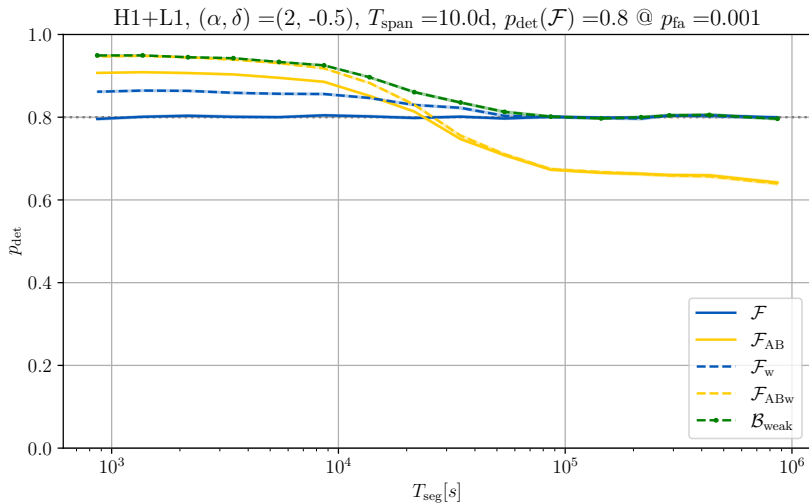
- 👉 generalized χ^2 -distribution with (2,2) degrees of freedom
- 👉 already includes segment weights $(\gamma A, \gamma B)_\ell$?

Compare:

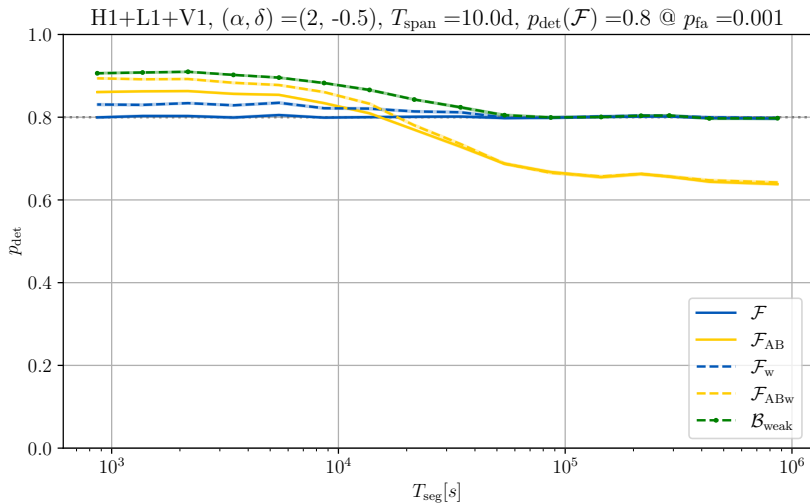
- $2\mathcal{F}(x) = 2D^{-1} [AB(\mathcal{F}_A(x) + \mathcal{F}_B(x)) - 2C^2\mathcal{F}_C(x)] \sim \chi_{4\text{dof}}^2$
- short-segment $\mathcal{F}_{AB}(x) = \begin{cases} \mathcal{F}_A(x) & \text{if } A > B \\ \mathcal{F}_B(x) & \text{otherwise} \end{cases} \sim \chi_{2\text{dof}}^2$



Results (H1+L1)



Results (H1+L1+V1)



- Half-normal h_0 -prior generalizes \mathcal{B} -statistic $\Rightarrow \mathcal{B}(x; H)$
- “weak-signal” limit $H \ll 1/\sqrt{\gamma} \Rightarrow$ new analytical solution!
- for longer segments: similar to \mathcal{F}_w , weaker than \mathcal{B}
- for short segments $\lesssim 1d$: equal or better than \mathcal{F}_{ABw}
- $\ln \mathcal{B}_{\text{weak}} = \frac{H^2}{10} \left(\underbrace{\sum_{\mu} x_{\mu}^2}_{\text{5-vector statistic}} - \text{Tr} \mathcal{M} \right)$ vs $2\mathcal{F} = x_{\mu} (\mathcal{M}^{-1})^{\mu\nu} x_{\nu}$
 \Rightarrow “5-vector statistic” [Astone+ CQG27(2010)], [D’Onofrio+ arXiv:2406.09236v]
- future work: better line-robust statistics $B_{S/GL}(H)$ based on proper h_0 -prior?