

# The weak-signal Bayes-factor approximation

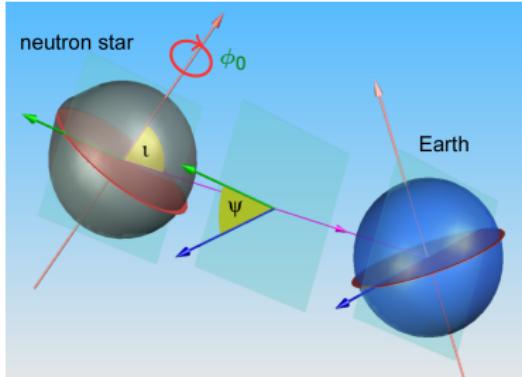
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Continuous gravitational waves and neutron stars @ Hannover  
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# Signal amplitude parameters



☞ Amplitude parameters:

$$\mathcal{A} = \{h_0, \underbrace{\cos \iota, \psi, \phi_0}_{?}\}$$

“radius” ( $h_0 \geq 0$ ) + 3 Euler angles

$\mathcal{A}$  generally unknown

Any spin orientation equally likely (isotropic):

$$P(\cos \iota | I) = \text{unif}(-1, 1)$$

$$P(\psi | I) = \text{unif}(0, 2\pi)$$

Any initial phase equally likely:

$$P(\phi_0 | I) = \text{unif}(0, 2\pi)$$

☞ “physical prior”  $P(h_0, \cos \iota, \psi, \phi_0 | I) = \text{const} \times \underbrace{P(h_0 | I)}_{?}$

## Part I: Pre-history

From  $\mathcal{F}(x)$ - to  $\mathcal{B}(x)$ -statistic

# JKS coordinates $\{\mathcal{A}^\mu\}$

[Jaranowski,Królak,Schutz PRD58(1998)]

“cartesian” JKS coordinates  $\mathcal{A}^\mu = \mathcal{A}^\mu(h_0, \cos \iota, \psi, \phi_0)$

$$\mathcal{A}^1 = 0.5h_0(1 + \cos^2 \iota) \cos 2\psi \cos \phi_0 - h_0 \cos \iota \sin 2\psi \sin \phi_0$$

$$\mathcal{A}^2 = 0.5h_0(1 + \cos^2 \iota) \sin 2\psi \cos \phi_0 + h_0 \cos \iota \cos 2\psi \sin \phi_0$$

$$\mathcal{A}^3 = -0.5h_0(1 + \cos^2 \iota) \cos 2\psi \sin \phi_0 - h_0 \cos \iota \sin 2\psi \cos \phi_0$$

$$\mathcal{A}^4 = -0.5h_0(1 + \cos^2 \iota) \sin 2\psi \sin \phi_0 + h_0 \cos \iota \cos 2\psi \cos \phi_0$$

## 👉 Signal strain @ detector

$$h(t; \mathcal{A}; \lambda) = \sum_{\mu=1}^4 \mathcal{A}^\mu h_\mu(t; \lambda)$$

$$h_1(t; \lambda) = a(t; \vec{n}_{\text{sky}}) \cos \phi(t; \lambda), \quad h_2(t; \lambda) = b(t; \vec{n}_{\text{sky}}) \cos \phi(t; \lambda)$$

$$h_3(t; \lambda) = a(t; \vec{n}_{\text{sky}}) \sin \phi(t; \lambda), \quad h_4(t; \lambda) = b(t; \vec{n}_{\text{sky}}) \sin \phi(t; \lambda)$$

Phase-evolution parameters:  $\lambda = \{\vec{n}_{\text{sky}}, \text{freq}, \text{spindown}, \text{binary}, \dots\}$

(here:  $\lambda$  assumed known)

# Likelihood ratio $\mathcal{L}(x; \mathcal{A}, \lambda)$

$\mathcal{H}_G$ : data = Gaussian noise:  $x(t) = n(t)$

$\mathcal{H}_S$ : data = Gaussian noise + signal:  $x(t) = n(t) + h(t; \mathcal{A}, \lambda)$

Goal: given  $x(t)$  "optimally" decide between  $\mathcal{H}_G$  and  $\mathcal{H}_S$

☞ likelihood ratio: Gaussian in  $\mathcal{A}^\mu$

$$\mathcal{L}(x; \mathcal{A}, \lambda) \equiv \frac{P(x|\mathcal{H}_S)}{P(x|\mathcal{H}_G)} = \exp \left[ -\frac{1}{2} \mathcal{A}^\mu \mathcal{M}_{\mu\nu} \mathcal{A}^\nu + \mathcal{A}^\mu x_\mu \right]$$

4 matched-filter outputs  $x_\mu(\lambda) \equiv (x|h_\mu(\lambda)) \in \mathbb{R}$

4x4 antenna-pattern matrix:

$$\mathcal{M}_{\mu\nu}(\vec{n}_{\text{sky}}) \equiv (h_\mu | h_\nu) = \underbrace{(S_n^{-1} T)}_{\gamma \equiv \text{"quality x quantity"}} \begin{pmatrix} A & C & O & \\ C & B & O & \\ O & O & A & C \\ C & B & C & B \end{pmatrix}$$

matched-filter scalar product:  $(x|y) \equiv S_n^{-1} \int x(t) y(t) dt = \gamma \langle xy \rangle$

$$A \equiv \langle a^2 \rangle, B \equiv \langle b^2 \rangle, C \equiv \langle ab \rangle$$

# 1998: Maximum-likelihood: $\mathcal{F}$ -statistic

☞ analytically maximize  $\mathcal{L}$  over  $\mathcal{A}^\mu$ : [Jaranowski,Królak,Schutz PRD58(1998)]

$$\mathcal{F}(x) \equiv \max_{\{\mathcal{A}^\mu\}} \ln \mathcal{L}(x; \mathcal{A}^\mu) = \frac{1}{2} x_\mu \left( \mathcal{M}^{-1} \right)^{\mu\nu} x_\nu$$

$$2\mathcal{F}(x) = 2D^{-1} \left[ AB(\mathcal{F}_A(x) + \mathcal{F}_B(x)) - 2C^2\mathcal{F}_C(x) \right]$$

$$2\mathcal{F}_A(x) \equiv \frac{x_1^2 + x_3^2}{\gamma A}, \quad 2\mathcal{F}_B(x) \equiv \frac{x_2^2 + x_4^2}{\gamma B}, \quad 2\mathcal{F}_C(x) \equiv \frac{x_1x_2 + x_3x_4}{\gamma C}$$

$$\text{sub-determinant } D = \det \begin{pmatrix} A & C \\ C & B \end{pmatrix} = AB - C^2$$

☞ Long thought to be "optimal", until ...

Neyman-Pearson-Searle theorem [Searle arXiv:0804.1161]

☞ Bayes factor is optimal statistic (if drawing from prior)

$$B_{S/G}(x) \equiv \int \mathcal{L}(x; \mathcal{A}) P(\mathcal{A}|I) d^4 \mathcal{A}$$

- “ $\mathcal{F}$ -statistic prior”  $P(\mathcal{A}^\mu) = \text{const}$  ☞  $B_{S/G} \propto e^{\mathcal{F}}$

(unphysical prior:  $P(h_0, \cos \iota, \psi, \phi_0) \propto h_0^3 (1 - \cos^2 \iota)^3$ )

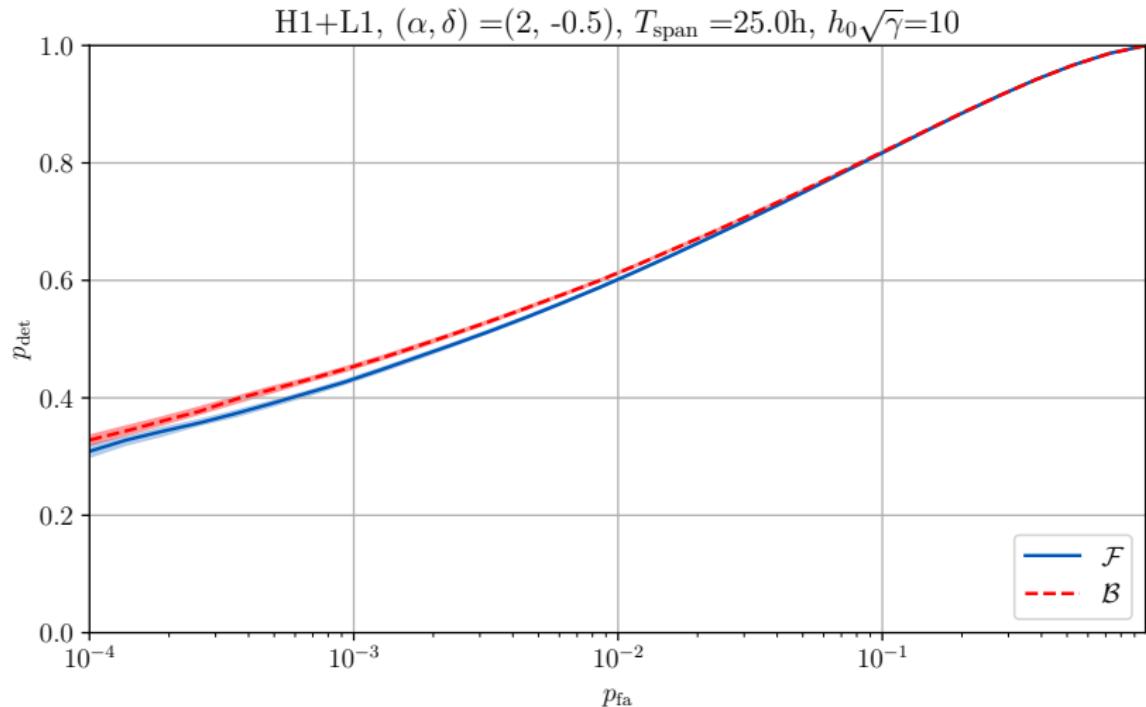
- “physical prior”  $P(h_0, \cos \iota, \psi, \phi_0) = \text{const} \times P(h_0)$ 
  - ☞ integrate  $\phi_0$  ✓
  - ☞ if  $P(h_0) = \text{const} \Rightarrow$  integrate  $h_0!$  ✓

“ $\mathcal{B}$ -statistic” ☞ 2D integral [Prix,Krishnan CQG26(2009)], [Whelan+ CQG31(2014)]

$$\mathcal{B}(x) \propto \int \frac{1}{\varrho} e^\Theta I_0(\Theta) d\cos \iota d\psi$$

$$\Theta(x; \cos \iota, \psi) \equiv \frac{q^2}{4\varrho^2}, q = q(x; \cos \iota, \psi), \varrho = \varrho(\cos \iota, \psi)$$

$\mathcal{B}(x)$  beats  $\mathcal{F}(x)$

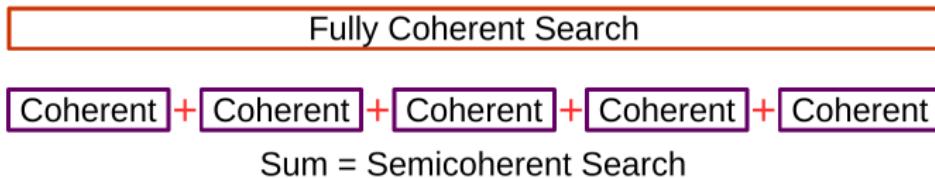


various attempts to solve/approximate  $\mathcal{B}(x)$ : [Dergachev PRD85(2012)][Whelan+ CQG31(2014)][Haris,Pai  
PRD96(2017)][Dhurandhar+ arXiv:1707.08163][Bero,Whelan CQG36(2019)][Wette Universe7(2021)]

## Part II: Recent surprises

### Short segments and weights

# Surprise 1: dominant-response $\mathcal{F}_{AB}$ -statistic [Covas,Prix PRD105(2022)]

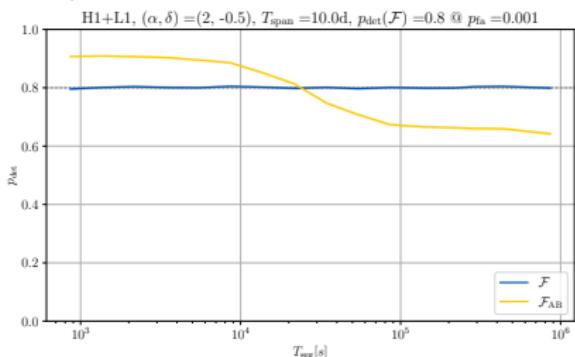


All-sky binary search ↗ very short segments  $T_{\text{seg}} \sim \mathcal{O}(100\text{s})$   
↗ degenerate  $\mathcal{F}(x)$  in certain  $\vec{n}_{\text{sky}}$  where  $D \rightarrow 0$ , ↗  
maximum-likelihood  $\mathcal{A}^{\mu}_{\text{ML}} = \mathcal{M}^{-1\mu\nu} x_{\mu}$  fails!

Fix: “dominant-response” statistic:

$$\mathcal{F}_{AB}(x) \equiv \begin{cases} \mathcal{F}_A(x) & \text{if } A > B \\ \mathcal{F}_B(x) & \text{otherwise} \end{cases}$$

↗ up to  $\sim 19\%$  more sensitive  
than  $\mathcal{F}(x)$  for short segments!!



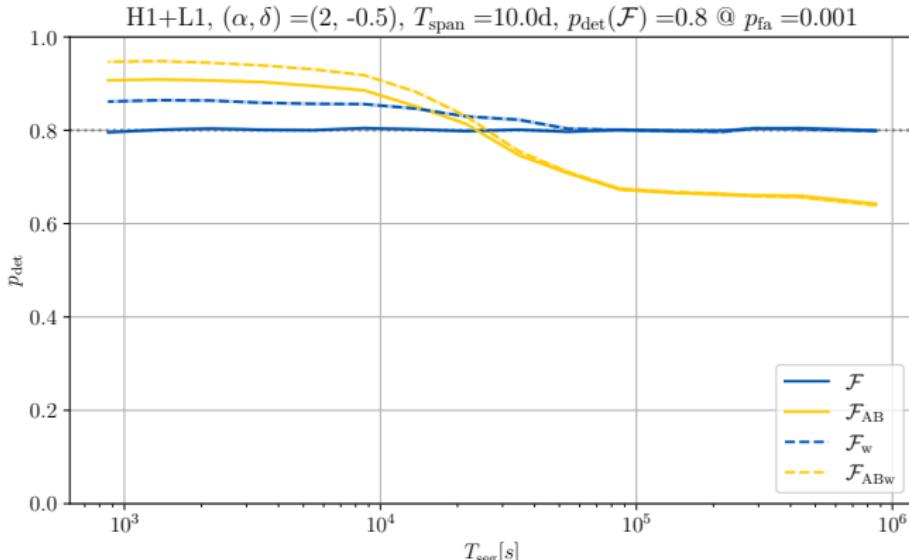
# Surprise 2: segment weighting

[Covas,Prix PRD106(2022)]

Can improve statistics by (empirical) *per-segment weights*  $w_\ell$ :

("data factor"  $\gamma = s_n^{-1} T$ )

- $\hat{\mathcal{F}}_w \equiv \sum_{\ell=1}^{N_{\text{seg}}} (\gamma(A + B))_\ell \mathcal{F}_\ell$
- $\hat{\mathcal{F}}_{ABw} \equiv \sum_{\ell=1}^{N_{\text{seg}}} \left( \gamma(Q + \frac{C^2}{Q}) \right)_\ell \mathcal{F}_{AB,\ell}$  (with  $Q = \max[A, B]$ )



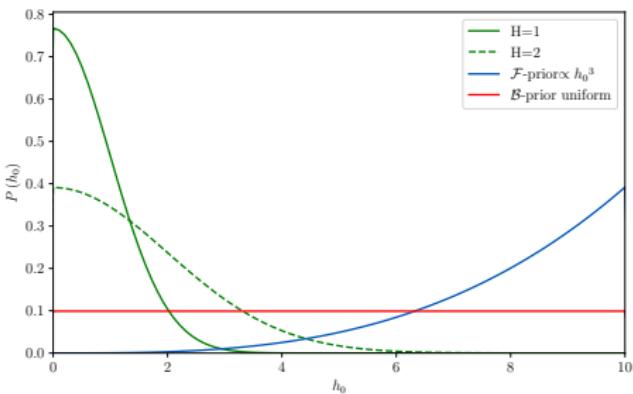
☞  $\mathcal{F}$ -statistic far from optimal for short segments!

### Part III Bayesian

Is there a ~~better~~ way?

Half-normal distribution with  $h_0$ -“scale”  $H$ 

$$P(h_0|I, H) = \frac{\sqrt{2}}{\sqrt{\pi}H} e^{-\frac{h_0^2}{2H^2}} \quad \text{for } h_0 \geq 0$$



- ✓ favors weak signals over strong
- ✓ proper prior
- ✓  $H \rightarrow \infty$  ↗ reduces to uniform- $h_0$
- ✓ conjugate to  $\mathcal{L}$ : can integrate  $h_0$ !

“relative scale”  $H_* \equiv H\sqrt{\gamma}$

$$\mathcal{B}(x; H_*) = \frac{1}{\pi} \int d\cos \iota \, d\psi \frac{1}{\sqrt{1 + H_*^2 \varrho^2}} e^\Theta I_0(\Theta)$$

$$\Theta(x; \cos \iota, \psi) \equiv \frac{H_*^2 q^2(x; \cos \iota, \psi)}{4(1 + H_*^2 \varrho^2(\cos \iota, \psi))}$$

Two interesting limits:

- strong-signal limit:  $H_* \gg 1$  ↗ recover const- $h_0$   $\mathcal{B}(x)$
- weak-signal limit:  $H_* \ll 1$ : Taylor-expand in  $H_*$ !

(coherent sensitivity  $h_0|_{P_{\text{fa}}=1\%}^{P_{\text{det}}=90\%} \sim 11.4/\sqrt{\gamma}$  ↗  $h_0\sqrt{\gamma} \sim \mathcal{O}(10)$ )

# Analytic weak-signal limit $H_* \ll 1$

Taylor-expand  $\mathcal{B}(x; H_*)$  in  $\mathcal{O}(H_*)$ :

$$\mathcal{B}(x) = 1 + \frac{H_*^2}{4} \langle q^2 - 2\varrho^2 \rangle_{\cos \iota, \psi} + \mathcal{O}(H_*^4)$$

## Weak-signal limit Bayes factor $\mathcal{B}_{\text{weak}}(x)$

$$\ln \mathcal{B}_{\text{weak}}(x) \equiv \frac{H^2}{10} \gamma [A(2\mathcal{F}_A(x) - 2) + B(2\mathcal{F}_B(x) - 2)]$$

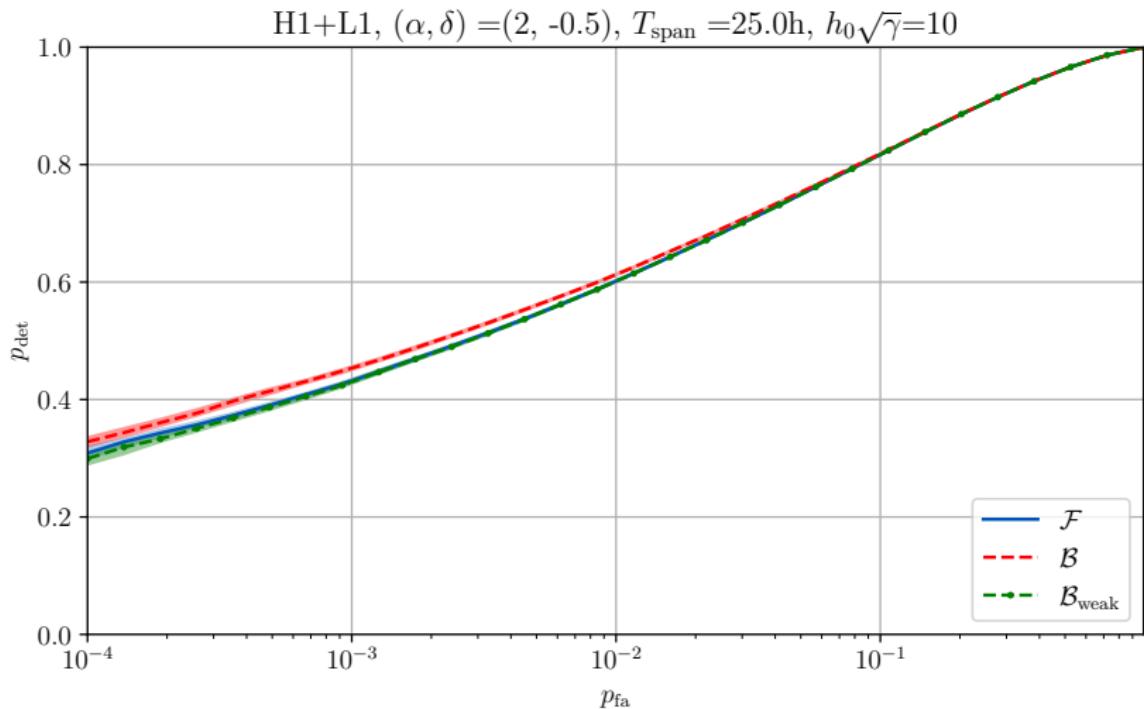
- generalized  $\chi^2$ -distribution with (2,2) degrees of freedom
- includes segment weights  $(\gamma A, \gamma B)_\ell$  for free?

Compare:

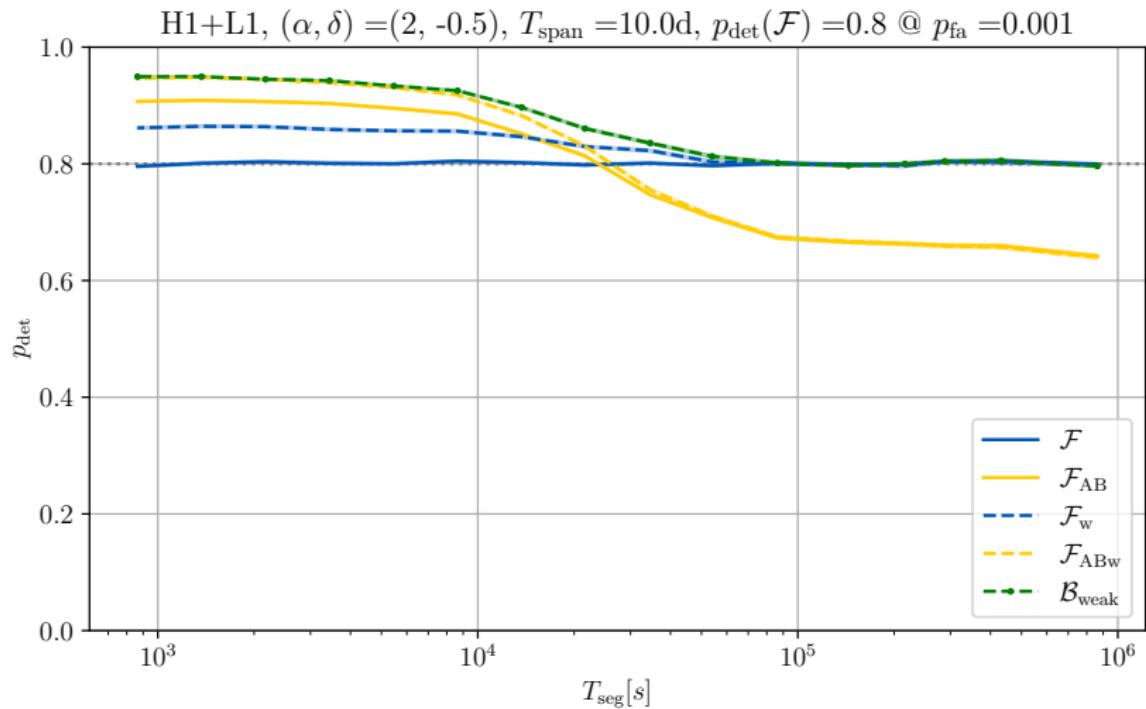
- $2\mathcal{F}(x) = 2D^{-1} [AB(\mathcal{F}_A(x) + \mathcal{F}_B(x)) - 2C^2\mathcal{F}_C(x)] \sim \chi^2_{4\text{dof}}$

- short-segment  $\mathcal{F}_{AB}(x) = \begin{cases} \mathcal{F}_A(x) & \text{if } A > B \\ \mathcal{F}_B(x) & \text{otherwise} \end{cases} \sim \chi^2_{2\text{dof}}$

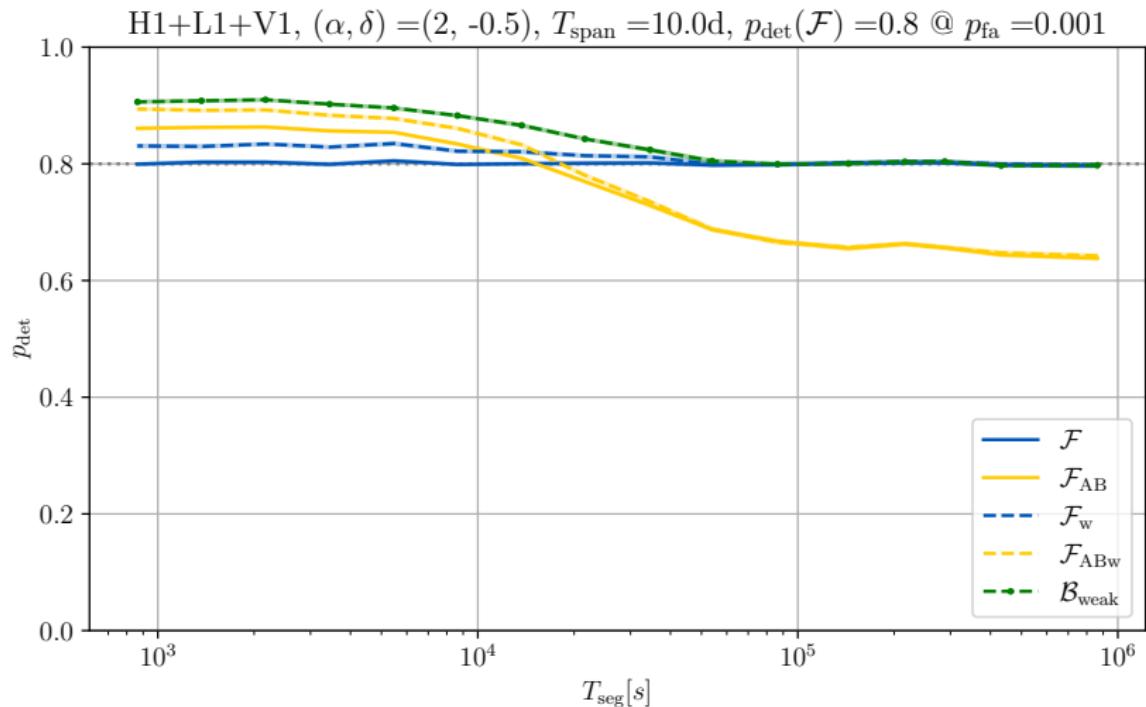
# Results



# Results (H1+L1)



# Results (H1+L1+V1)



- Half-normal  $h_0$ -prior generalizes  $\mathcal{B}$ -statistic  $\rightarrow \mathcal{B}(x; H)$
- “weak-signal” limit  $H \ll 1/\sqrt{\gamma} \rightarrow$  new analytical solution
- for longer segments: similar to  $\mathcal{F}_w$ , weaker than  $\mathcal{B}$
- for short segments  $\lesssim 1d$ : similar/better than  $\mathcal{F}_{ABw}$
- $\ln \mathcal{B}_{\text{weak}} = \frac{H^2}{10} \left( \underbrace{\sum_{\mu} x_{\mu}^2}_{\text{---}} - \text{Tr} \mathcal{M} \right) \quad \text{vs} \quad 2\mathcal{F} = x_{\mu} (\mathcal{M}^{-1})^{\mu\nu} x_{\nu}$   
 $\rightarrow$  “5-vector statistic” [Astone+ CQG27(2010)], [D’Onofrio+ arXiv:2406.09236v]
- better line-robust statistics  $B_{S/GL}(H)$ ? (proper prior)